Predictions on the second-class current decays

 $au^- o \pi^- \eta^{(\prime)}
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Objectives

- 1. We reassess the $au^- o \pi^- \eta^{(\prime)}
 u_ au$ decays, unseen in Nature so far.
- 2. To motivate near future B-factories e.g. Belle-II, their discovery.

Introduction

- au $au^- au^- \eta^{(\prime)}
 u_ au$ decays belong to the so-called second-class current processes: parity conservation implies that these decays must proceed through the vector current, which has opposed G-parity to the $\pi^-\eta^{(\prime)}$ system.
- In the limit of exact isospin symmetry G-parity is exact and these processes are forbidden.
- Isospin is however an approximate symmetry, slightly broken both by $m_u \neq m_d$ (in QCD) and $q_u \neq q_d$ (in QED), which results in a sizable suppression of these decays.

Decay width

Amplitude of the process

$$\mathcal{M} = rac{G_F}{\sqrt{2}} V_{ud} \bar{u}(
ho_{
u_{ au}}) \gamma_{\mu} (1-\gamma_5) u(
ho_{ au}) \langle \pi^- \eta^{(\prime)} | \bar{d} \gamma^{\mu} u | 0 \rangle$$
.

▶ Hadronic matrix element in terms of form factors

$$\langle \pi^{-} \eta^{(\prime)} | \bar{d} \gamma^{\mu}_{u} | \mathbf{0} \rangle =$$

$$\left[(\rho_{\eta^{(\prime)}} - \rho_{\pi})^{\mu} + \frac{\Delta_{\pi^{-} \eta^{(\prime)}}}{s} q^{\mu} \right] c_{\pi \eta^{(\prime)}}^{V} F_{+}^{\pi \eta^{(\prime)}}(s) + \frac{\Delta_{K^{0}K^{+}}^{QCD}}{s} q^{\mu} c_{\pi^{-} \eta^{(\prime)}}^{S} F_{0}^{\pi^{-} \eta^{(\prime)}}(s) .$$

▶ The finiteness of the matrix element at the origin imposes

$$F_{+}^{\pi^{-}\eta^{(\prime)}}(\mathbf{0}) = -\frac{c_{\pi^{-}\eta^{(\prime)}}^{S} \Delta_{K^{0}K^{+}}^{QCD}}{c_{\pi^{-}\eta^{(\prime)}}^{V} \Delta_{\pi^{-}\eta^{(\prime)}}^{QCD}} F_{0}^{\pi^{-}\eta^{(\prime)}}(\mathbf{0}), \quad \Delta_{PQ} = m_{P}^{2} - m_{Q}^{2}. \quad (1)$$

 \blacktriangleright Differential decay width as a function of the $\pi^-\eta^{(\prime)}$ invariant mass

$$\frac{d\Gamma\left(\tau^{-} \to \pi^{-} \eta^{(\prime)} \nu_{\tau}\right)}{d\sqrt{s}} = \frac{G_{F}^{2} M_{\tau}^{3}}{24\pi^{3} s} S_{EW} |V_{ud} F_{+}^{\pi^{-} \eta^{(\prime)}}(\mathbf{0})|^{2} \left(1 - \frac{s}{M_{\tau}^{2}}\right)^{2} \\
\left\{ \left(1 + \frac{2s}{M_{\tau}^{2}}\right) q_{\pi^{-} \eta^{(\prime)}}^{3}(s) |\widetilde{F}_{+}^{\pi^{-} \eta^{(\prime)}}(s)|^{2} + \frac{3\Delta_{\pi^{-} \eta^{(\prime)}}^{2}}{4s} q_{\pi^{-} \eta^{(\prime)}}(s) |\widetilde{F}_{0}^{\pi^{-} \eta^{(\prime)}}(s)|^{2} \right\},$$

with $q_{PQ}(s) = \frac{\sqrt{s^2 - 2s(m_P^2 + m_Q^2) + \Delta_{PQ}^2}}{2\sqrt{s}}$, and where $\widetilde{F}_{+,0}^{\pi^- \eta^{(\prime)}}(s) = \frac{F_{+,0}^{\pi^- \eta^{(\prime)}}(s)}{F_{+,0}^{\pi^- \eta^{(\prime)}}(0)}$, are the two form factors normalised to unity at the origin.

Form factors parameterizations

Vector Form Factor

we employ Chiral Perturbation Theory with Resonances [1] $F_{+}^{\pi^{-}\eta}(s) = \varepsilon_{\pi\eta^{(\prime)}} \left[1 + \sum_{V} \frac{F_{V}G_{V}}{F^{2}} \frac{s}{M_{V}^{2} - s} \right] = \varepsilon_{\pi\eta^{(\prime)}} F_{+}^{\pi^{-}\pi^{0}}(s).$

where $\varepsilon_{\pi\eta}$ and $\varepsilon_{\pi\eta'}$ accounts for the $\pi-\eta^{(\prime)}$ mixing.

- Scalar Form Factor
- Chiral Perturbation Theory with Resonances [1]

$$F_0^{\pi^-\eta^{(\prime)}}(s) = c_0^{\pi^-\eta^{(\prime)}} \frac{M_S^2 + \Delta_{\pi^-\eta^{(\prime)}}}{M_S^2 - s - iM_S \Gamma_S(s)}, c_0^{\pi^-\eta} = \cos\theta_{\eta\eta'} - \sqrt{2}\sin\theta_{\eta\eta'} \\ c_0^{\pi^-\eta'} = \cos\theta_{\eta\eta'} + \frac{\sin\theta_{\eta\eta'}}{\sqrt{2}},$$
(3)
From Eqs. (1,2,3) we find
$$\varepsilon_{\pi\eta} = (9.8 \pm 0.3) \cdot 10^{-3}$$

$$\varepsilon_{\pi\eta'} = (2.5 \pm 1.5) \cdot 10^{-4}$$
Dispersion Relations

Dispersion Relations

Elastic unitarity: Omnès equation

$$F_0^{\pi^-\eta^{(\prime)}}(s) = P(s) \exp \left[\frac{s - s_0}{\pi} \int_{s_{th}}^{\infty} ds' \frac{\delta_{1,0}^{\pi^-\eta^{(\prime)}}(s')}{(s' - s_0)(s' - s - i\varepsilon)} \right],$$

where $\delta_{1,0}^{\pi^-\eta^{(\prime)}}(s) = \arctan \frac{\operatorname{Im} t_{1,0}^{\pi^-\eta^{(\prime)}}(s)}{\operatorname{Re} t_{1,0}^{\pi^-\eta^{(\prime)}}(s)}$ with $t^{\pi^-\eta^{(\prime)}}(s)$ from Ref. [2].

▶ Inelastic cuts: $\pi^-\eta$ coupled to $\pi^-\eta'$ and/or $\kappa^-\kappa^0$ and viceversa [1].

Results: Branching ratio predictions

 $ightharpoonup au^- o \pi^- \eta
u_ au$

Scalar Form factor	Contribution (10^{-5})	Total (10 ⁻⁵)
ChPT (1 resonance)	$0.72^{+0.46}_{-0.22}$	0.98(51)
ChPT (2 resonances)	$0.48^{+0.29}_{-0.14}$	0.74(32)
Elastic approximation (Omnès)	$0.10^{+0.02}_{-0.03}$	0.36(4)
$\pi^-\eta$ coupled to $\pi^-\eta'$	0.15(9)	0.41(9)
$\pi^-\eta$ coupled to $\kappa^-\kappa^0$	1.86(11)	2.12(11)
$\pi^-\eta$ coupled to $\pi^-\eta'$ and $\kappa^-\kappa'$	1.41(9)	1.67(9)

Table 1: Total branching ratio includes the vector, $0.26(2) \cdot 10^{-5}$, and the scalar contributions.

Scalar Form factor	Contribution
ChPT (1 resonance)	$[2 \cdot 10^{-11}, 7 \cdot 10^{-10}]$
ChPT (2 resonances)	$[5 \cdot 10^{-11}, 2 \cdot 10^{-9}]$
Elastic approximation (Omnès)	$[2 \cdot 10^{-9}, 4 \cdot 10^{-8}]$
$\pi^-\eta$ coupled to $\pi^-\eta'$	$[2 \cdot 10^{-7}, 2 \cdot 10^{-6}]$
$\pi^-\eta$ coupled to $\kappa^-\kappa^0$	$[3 \cdot 10^{-7}, 3 \cdot 10^{-6}]$
$\pi^-\eta$ coupled to $\pi^-\eta'$ and $\kappa^-\kappa^0$	$[1 \cdot 10^{-7}, 1 \cdot 10^{-6}]$

Table 2: Scalar contribution to the branching ratio. The (subleading) vector gives $[0.3, 6] \cdot 10^{-11}$.

Results: Invariant mass distribution

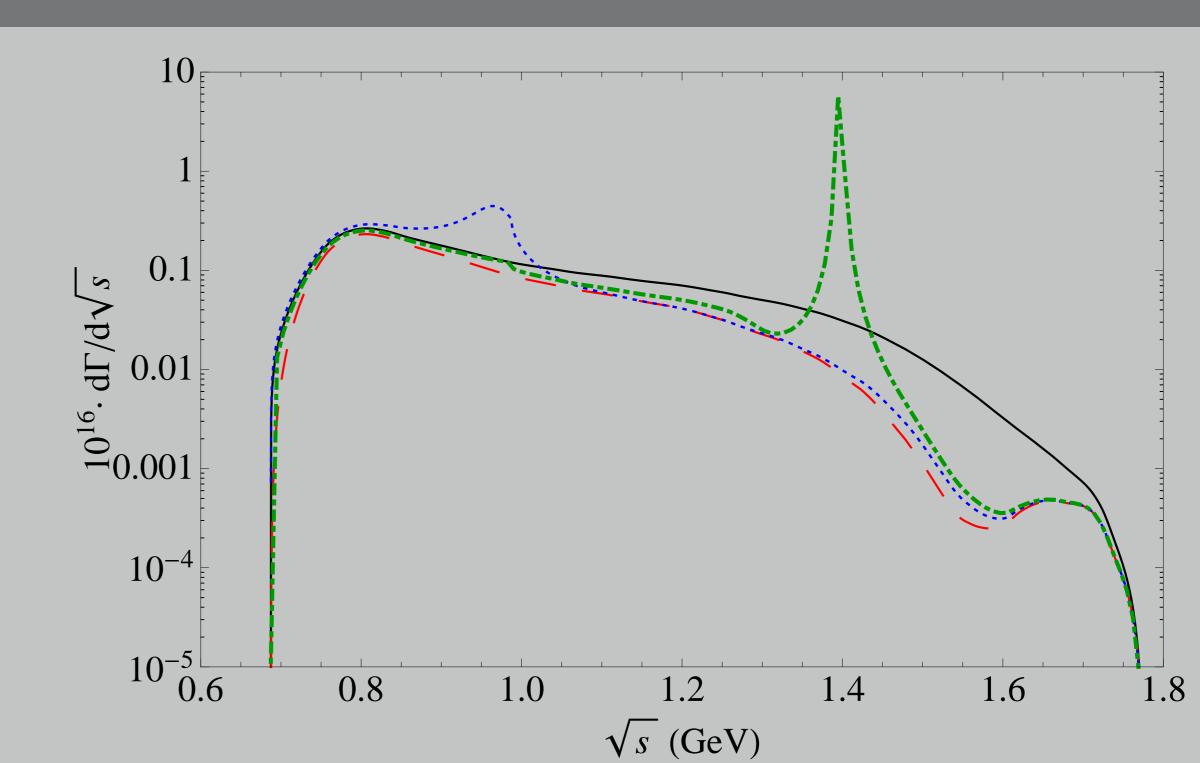


Figure 1: $au^- o \pi^- \eta
u_ au$ distribution. We show the vector contribution (red-dashed curve) and the full distribution as obtained by employing: the scalar form factor in its elastic version (black solid curve), the three coupled-channels analysis (green dot-dashed curve) and ChPT with two resonances (blue dotted curve).

Conclusions

- \blacktriangleright We focus on the Standard Model prediction of $au^- o \pi^- \eta^{(\prime)}
 u_ au$ decays.
- ► We have addressed the description of the participant vector and scalar form factors based on Chiral Perturbation Theory with resonances supplemented by dispersion relations.
- According to our results, the discovery of second-class currents might be possible at Belle-II thanks to the increased luminosity with respect to its predecessors.

References

- 1 R. Escribano, S. Gonzàlez-Solís and P. Roig, arXiv:1601.03989.
- 2 Z. H. Guo and J. A. Oller, Phys. Rev. D 84, 034005 (2011).

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