# Analysis of the production mechanism of $p+p \rightarrow p+K+\Lambda$ using PWA 

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## Outline

- Introduction - Strangeness Production
- Experimental Data
- Method of Partial Wave Analysis
- Results of Combined Analysis
- Excitation Function
- Scattering Length
- Summary
- Outlook


## Strangeness Production



Detailed Understanding required:

- Transport Model
- Production of Exotic


## One - Boson Exchange Modell



## Strangeness Production



## Strangeness Production

| Resonance | $J^{P}$ | Mass $\left(\mathrm{GeV} / \mathrm{c}^{2}\right)$ | $\Gamma\left(\mathrm{MeV} / \mathrm{c}^{2}\right)$ |
| :--- | :--- | :--- | :--- |
| $\mathrm{N}^{*}(1650)$ | $1 / 2^{-}$ | 1.655 | 0.150 |
| $\mathrm{~N}^{*}(1710)$ | $1 / 2^{+}$ | 1.710 | 0.100 |
| $\mathrm{~N}^{*}(1720)$ | $3 / 2^{+}$ | 1.720 | 0.250 |
| $\mathrm{~N}^{*}(1875)$ | $3 / 2^{-}$ | 1.875 | 0.220 |
| $\mathrm{~N}^{*}(1880)$ | $1 / 2^{+}$ | 1.870 | 0.235 |
| $\mathrm{~N}^{*}(1895)$ | $1 / 2^{-}$ | 2.090 | 0.090 |
| $\mathrm{~N}^{*}(1900)$ | $3 / 2^{+}$ | 1.900 | 0.0250 |



## Strangeness Production



Meson 2016 - Kraк.

## Strangeness Production

LO - NLO Calculations

J. Haidenbauer et al., Nucl. Phys. A915, 24 (2013)

## Strangeness Production


S.Abd El-Samad, Eur.Phys.J A49(2013)


At Threshold : $2130 \mathrm{MeVc}^{-2}$

## Strangeness Production



FINUDA

G. Agakishiev,Phys.Lett. B742 (2015) 242-248



# Strangeness Production 

# Strangeness Production 

## Experimental Data

## Data Sets

| Experiment | $E_{B}[\mathrm{GeV}]$ | Statistics | Reference |
| :---: | :---: | :---: | :---: |
| DISTO | 2.15 | 121k | M. Maggiora et al. Nucl Phys. A835, 43 (2010) M.Maggiora, Nucl Phys. A691, 329 (2001) |
| COSY-TOF | 2.16 | 43k | M. Roeder et al., Eur. Phys..J. A49, 157 (2013) S. Jowzaee et al., Eur. Phys. J. A52, 7 (2016) |
| DISTO | 2.5 | 304k | M. Maggiora et al. Nucl Phys. A835, 43 (2010) M.Maggiora, Nucl Phys. A691, 329 (2001) |
| DISTO | 2.85 | 424k | M. Maggiora et al. Nucl Phys. A835, 43 (2010) M.Maggiora, Nucl Phys. A691, 329 (2001) <br> F. Balestra et al. , Phys.Rev.Lett.83. 1534 (1999) |
| FOPI | 3.1 | 903 | R. Münzer, Hyp. Int., 233,1-3,159-166 (2015) |
| HADES | 3.5 | 21k | G. Agakishiev,Phys.Lett. 8742 (2015) 242-248 |

## The FOPI Experiment

## SIS18 GSI Darmstadt



## Beam Energy: 3.1 GeV

- Fixed-target Setup
- Full azimuthal coverage, $5^{\circ}-110^{\circ}$ in polar angle
- Momentum resolution $\approx 7 \%-15 \%$
- Particle identification via $\mathrm{dE} / \mathrm{dx}$ \& ToF

Trigger Detector - $\mathrm{Si} \wedge \mathrm{ViO}$ :
$\Lambda$ - Enhancement: $\quad 14.1 \pm 7.9(\text { stat })_{-0.6}^{+4.3}$

## The HADES experiment

High Acceptance Di-electron Spectrometer GSI, Darmstadt


## Beam Energy: 3.5 GeV

- Fixed-target Setup
- Full azimuthal coverage, $15^{\circ}-185^{\circ}$ in polar angle
- Momentum resolution $\approx 1 \%-5 \%$
- Particle identification via dE/dx \& ToF

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HADES Coll. (G. Agakishiev et al.),
Eur. Phys. J. A41 (2009)
```


## COSY-TOF Spectrometer



Acceptance: $1^{\circ}-60^{\circ}$ (polar), $2 \pi$ (azimuthal)


Sec. Vertex: $\sigma_{x, y}<0.5 \mathrm{~mm}, \sigma_{z}<2.5 \mathrm{~mm}$
Momentum resolution $\approx 1 \%-1.5 \%$

$$
\sigma_{M M(p K)}=1.5 \mathrm{MeV} / \mathrm{c}^{2}
$$

## DISTO Spectrometer



Acceptance: $23^{\circ}-43^{\circ}$ (polar), $2 \pi$ (azimuthal)

$$
\begin{gathered}
\sigma_{p}=5 \% \\
\sigma_{M M(p K)}=30 \mathrm{MeV} / \mathrm{c}^{2}
\end{gathered}
$$

## Partial Wave Analysis

## Bonn-Gatchina PWA Framework

A. Sarantsev et.al., Eur. Phys J A 252005

Cross-section Decomposition

$$
d \sigma=\frac{(2 \pi)^{4}|A|^{2}}{4|k| \sqrt{s}} d \phi\left(P, q_{1}, q_{2}, q_{3}\right), \quad P=k_{1}+k_{2}
$$

$A$ : reaction amplitude $\mathrm{A} \propto A \propto A_{t r}^{\alpha}(\mathrm{s}) \quad$ (Transition amplitude of wave $\alpha$ )
$k: 3$-momentum of the initial particle in the CM
$s-P^{2}:\left(k_{1}+k_{2}\right)^{2}$
$d \phi\left(P, q_{1}, q_{2}, q_{3}\right)$ : invariant three-particle phase space

Parameterization of the Transition

$$
\begin{array}{ll}
a_{1}^{\alpha} & \text { Constant amplitude } \\
a_{2}^{\alpha} & \text { Phase } \\
a_{3}^{\alpha} & \text { Energy dependent amp. }
\end{array}
$$

## Why PWA?

## Masses


(a)

(b)


(e)
(d)

(g)

(h)
Hel. - Angle

(j)

(i)
(f)

(k)


+ Experimental Data

$\mathrm{pp} \rightarrow \mathrm{p} \mathrm{K}{ }^{+} \wedge$ Phase Space


## Why PWA?



## The $\Sigma \mathrm{N}$ Cusp Effect

## Coupled Channel:



At Threshold : $2130 \mathrm{MeVc}^{-2}$

Quantum Number of Cusp: $0^{+} / 1^{+}(\mathrm{L}=0,2)$

S.Abd El-Samad, Eur.Phys.J A49(2013)

## Cusp Spectral Function

The Breit-Wigner:

$$
\frac{d \sigma_{p \Lambda}}{d m_{p \Lambda}} \approx \frac{1}{\left|m_{R}^{2}-m_{p \Lambda}^{2}-i m_{p \Lambda} \Gamma\right|^{2}}
$$

Mass $M_{\text {cusp }}=2.13 \mathrm{GeV}$, With $\Gamma=0.02 \mathrm{GeV}$

## Combined Analysis

## Parameter Scan

## Variation of Included N* Resonances <br> Five best solution used to obtain systematical error

| Solution | A | B | C | D | E |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Loglike | -67142 | -67018 | -66878 | -66405 | -66393 |
| $N^{*}(1650)$ | + | + | + | + | + |
| $N^{*}(1710)$ | + | + | + | + | + |
| $N^{*}(1720)$ | + | + | + | - | - |
| $N^{*}(1875)$ | + | + | - | + | + |
| $N^{*}(1880)$ | + | + | + | + | + |
| $N^{*}(1895)$ | + | + | + | + | + |
| $N^{*}(1900)$ | - | + | + | + | + |
| $\Sigma N-C u s p$ | + | + | + | + | + |

## PWA Results



## PWA Results

- Comparison of Experimental Data and Results of PWA
- Using 3 Mass and 9 Angular Spectra

| Experi <br> ment | $\mathrm{E}_{\mathrm{B}}[\mathrm{GeV}]$ | $\mathrm{X}^{2} / \mathrm{ndf}$ |
| :---: | :---: | :---: |
| DISTO | 2.15 | 1.52 |
| COSY- <br> TOF | 2.16 | 0.44 |
| DISTO | 2.5 | 2.56 |
| DISTO | 2.85 | 3.55 |
| FOPI | 3.1 | 0.91 |
| HADES | 3.5 | 2.14 |

## Total Cross Section


R.Muenzer et al.,Hyperfine 233, 159-166 (2016)

Value:

$$
\sigma_{p K \Lambda}=C_{1}\left(1-\frac{s_{0}}{\left(\sqrt{s_{0}}+\epsilon\right)^{2}}\right)^{C_{2}}\left(\frac{s_{0}}{\left(\sqrt{s_{0}}+\epsilon\right)^{2}}\right)^{C_{3}} \quad \begin{aligned}
& C_{1}=1.18 \pm 0.4210^{3} \\
& C_{2}=1.88 \pm 0.12 \\
& C_{3}=3.28 \pm 0.71
\end{aligned}
$$

A. Sibirtsev et.al. Nucl.Phys. A632,131 (1998)

## Excitation Function

$$
\mathrm{P}+\mathrm{p} \rightarrow \mathrm{p}+\mathrm{K}^{+}+\Lambda \text { (non resonant) }
$$



Combined data set of data samples, provide required constraint on contributions

## Cross Section

$$
P+p \rightarrow p+N^{*}\left(->K^{+}+\Lambda\right)
$$






## Final State Interaction in PWA

$$
A_{2 b}^{\beta}=\frac{\sqrt{s_{i}}}{1+\frac{1}{2} r^{\beta} q^{2} a_{p \Lambda}^{\beta}+i q a_{p \Lambda}^{\beta} q^{2 L} / F\left(q, r^{\beta}, L\right)}
$$

$a_{p \Lambda}^{\beta}$ Scattering Length
$r^{\beta}$ Effective Range of System

$$
\begin{array}{cc}
\alpha_{s}=-1.43 \pm 0.36 \pm 0,09 \mathrm{fm} & \alpha_{t}=-1.88 \pm 0.38 \pm 0.10 \mathrm{fm} \\
r_{s}=1.31 \pm 0.24 \pm 0,16 \mathrm{fm} & r_{t}=1.04 \pm 0.78 \pm 0.15 \mathrm{fm}
\end{array}
$$

| Source | ${ }^{1} S_{0} \alpha_{\Lambda-p}$ | ${ }^{1} S_{0} r_{\Lambda-p}$ | ${ }^{3} S_{1} \alpha_{\Lambda-p}$ | ${ }^{3} S_{1} r_{\Lambda-p}$ | $\left\langle\alpha_{\Lambda-p}\right\rangle$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| NLO[1] | -2.91 fm | 2.78 fm | -1.54 fm | 2.72 fm | -1.88 fm |
| LO[2] | -1.91 fm | 1.40 fm | -1.23 fm | 2.13 fm | -1.4 fm |
| $[2]$ | $-1.8_{-4.2}^{+2.3} \mathrm{fm}$ |  | $-1.6_{-0.8}^{+1.1} \mathrm{fm}$ |  | $-1.25 \pm 0.08 \pm 0.03 \mathrm{fm}$ |
| COSY- <br> TOF[3] |  |  |  | $-1.31_{-0.49}^{+3.2} \pm 0.3$ <br> $\pm 0.16 \mathrm{fm}$ |  |
| COSY- <br> TOF[4] |  |  |  | $-1.233 \pm 0.014 \pm 0.3 \pm$ |  |

[1] Nucl.Phys. A915, 24(2013) [2] Phys.Rev. 173,1452 (1968)
[3] Eur. Phys.J. A49, 157 (2013) / [4] F.Hauenstein, PhD Thesis (2014)

## Summary

- Strangeness production has to understood in elementary reaction
- Different contribution Channels ( $\mathrm{N}^{*}$, Cusp Effect, FSI)
- Description requires Partial Wave Analysis.

Allow to analyze different beam data in parallel.
Combined analysis necessary to provide sufficient information

- Extraction of Excitation Function of N*
- $\mathrm{N}^{*}$ play dominate role in the production at GeV Energies
- Scattering Length and effective extraction
- Values for Singlet and Triplet separately
- Separation between LO and NLO difficult.


## Outlook - $\Sigma$-N Cusp

The Breit-Wigner:

$$
\frac{d \sigma_{p \Lambda}}{d m_{p \Lambda}} \approx \frac{1}{\left|m_{R}^{2}-m_{p \Lambda}^{2}-i m_{p \Lambda} \Gamma\right|^{2}}
$$



Above the threshold

$$
\begin{gathered}
g_{p \Sigma} \ll g_{p \Lambda} \text { Symmetric } \\
g_{p \Sigma} \gg g_{p \Lambda} \text { Antisymmetric }
\end{gathered}
$$

Below the threshold

The Flatté parametrization:

$$
\frac{d \sigma_{p \Lambda}}{d m_{p \Lambda}} \approx \frac{\Gamma_{p \Lambda}}{\left|m_{R}^{2}-m_{p \Lambda}^{2}-i m_{p \Lambda}\left(\Gamma_{p \Lambda}+\Gamma_{p \Sigma}\right)\right|^{2}}
$$

$$
\Gamma_{p \Lambda}=g_{p \Lambda} * q_{p \Lambda} \quad \Gamma_{p \Sigma}=g_{p \Sigma} * q_{p \Sigma}
$$

$$
q_{p \Sigma}=\frac{\sqrt{\left(m_{p \Sigma}^{2}-\left(m_{\Sigma}+m_{p}\right)^{2}\right) *\left(m_{p \Sigma}^{2}-\left(m_{p}-m_{\Sigma}\right)^{2}\right)}}{2 m_{p \Sigma}}
$$

$$
q_{p \Sigma}=i * \frac{\sqrt{\left(\left(m_{\Sigma}+m_{p}\right)^{2}-m_{p \Sigma}^{2}\right) *\left(m_{p \Sigma}^{2}-\left(m_{p}-m_{\Sigma}\right)^{2}\right)}}{2 m_{p \Sigma}}
$$

S. Jowzaee et al., Eur. Phys. J. A52, 7 (2016)

Incoherent Analysis by COSY-TOF

## Outlook - $\Sigma$-N Cusp : Combined Analysis

| Experiment | $\mathrm{E}_{\mathrm{B}}[\mathrm{GeV}]$ | Statistics |
| :---: | :---: | :---: |
| DISTO | 2.15 | 121 k |
| COSY-TOF | 2.16 | 43 k |
| DISTO | 2.5 | 304 k |
| DISTO | 2.85 | 424 k |



Resulting value:

$$
\begin{gathered}
g_{p \Lambda}=0.38 \pm 0.06 \pm 0.008 \quad 10^{-2} \\
g_{p \Sigma}=1.60 \pm 0.07 \pm 0.03 \quad 10^{-2}
\end{gathered}
$$

## Thank You

## HADES Collaboration

DISTO Collaboration
M. Maggiora

## FOPI Collaboration

## COSY-TOF Collaboration

 J. Ritman, E. Roderburg F. Hauenstein, D. Gronzka

K-Cluster - Excellence Cluster Universe - TU Munich
L. Fabbietti, E. Epple, P. Klose, S.Lu, J. Siebenson, D. Soliman

## Thank for your attention

## Final State Interaction






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## Branching Ratio

|  | Mass $\left[\mathrm{GeV} / \mathrm{c}^{2}\right]$ | Width $\left[\mathrm{GeV} / \mathrm{c}^{2}\right]$ | $\Gamma_{\Lambda K} / \Gamma_{\text {All }} \%$ |
| :--- | :---: | :---: | :---: |
| $\mathrm{~N}(1650) \mathrm{S}_{11}$ | 1.655 | 0.150 | $3-11$ |
| $\mathrm{~N}(1710) \mathrm{P}_{11}$ | 1.710 | 0.200 | $5-25$ |
| $\mathrm{~N}(1720) \mathrm{D}_{13}$ | 1.720 | 0.250 | $1-15$ |
| $\mathrm{~N}(1875) \mathrm{D}_{13}$ | 1.875 | 0.220 | $4 \pm 2$ |
| $\mathrm{~N}(1880) \mathrm{P}_{11}$ | 1.870 | 0.235 | $2 \pm 1$ |
| $\mathrm{~N}(1895) \mathrm{S}_{11}$ | 1.895 | 0.090 | $18 \pm 5$ |
| $\mathrm{~N}(1900) \mathrm{P}_{13}$ | 1.900 | 0.250 | $0-10$ |

