

# Electromagnetic effects on meson production: a new tool for studying the space-time evolution of heavy ion collisions.



Andrzej Rybicki

H. Niewodniczański Institute of Nuclear Physics  
Polish Academy of Sciences

- 1) Introduction ;
- 2) EM effects in heavy ion collisions ;
- 3) Space-time evolution of the system ;
- 4) Summary & outlook.

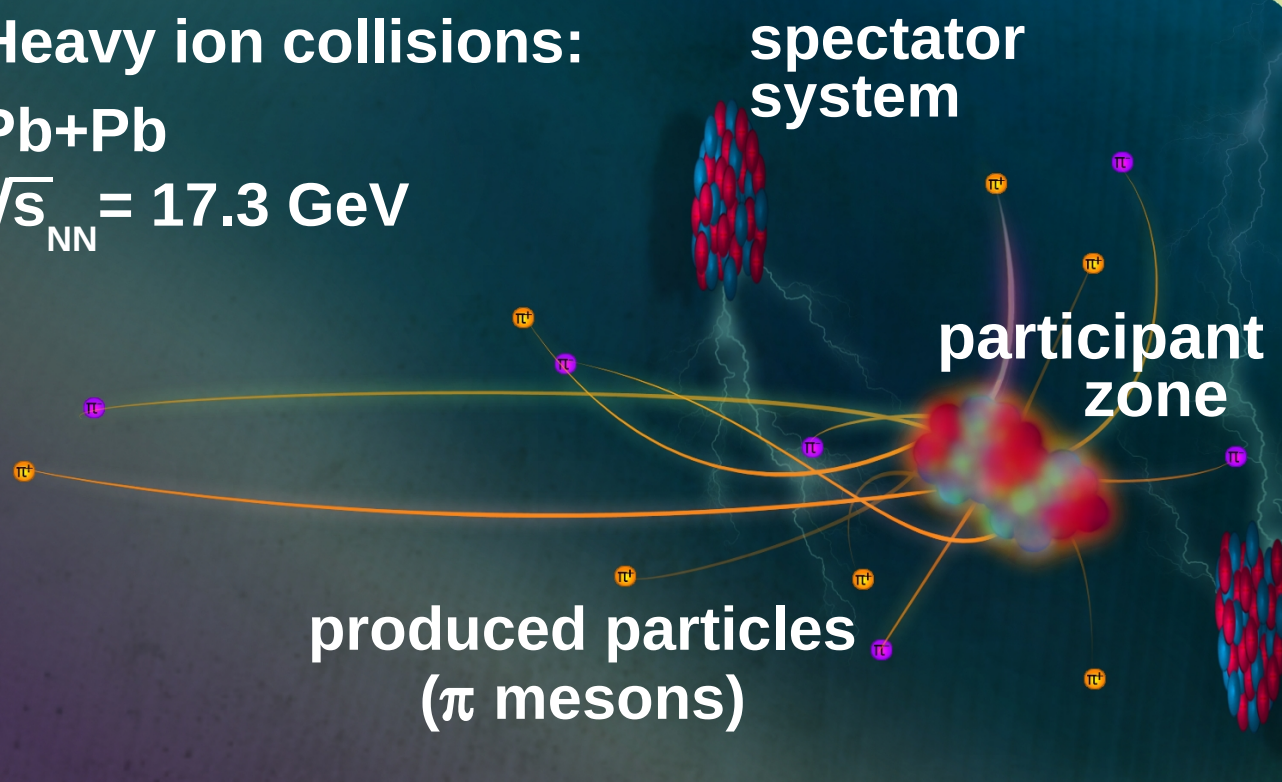
work in collaboration with  
Antoni Szczurek  
Mariola Kłusek-Gawenda  
Nikolaos Davis  
Vitalii Ozvenchuk  
Mirosław Kiełbowski

# ***1) Introduction***

Heavy ion collisions:

Pb+Pb

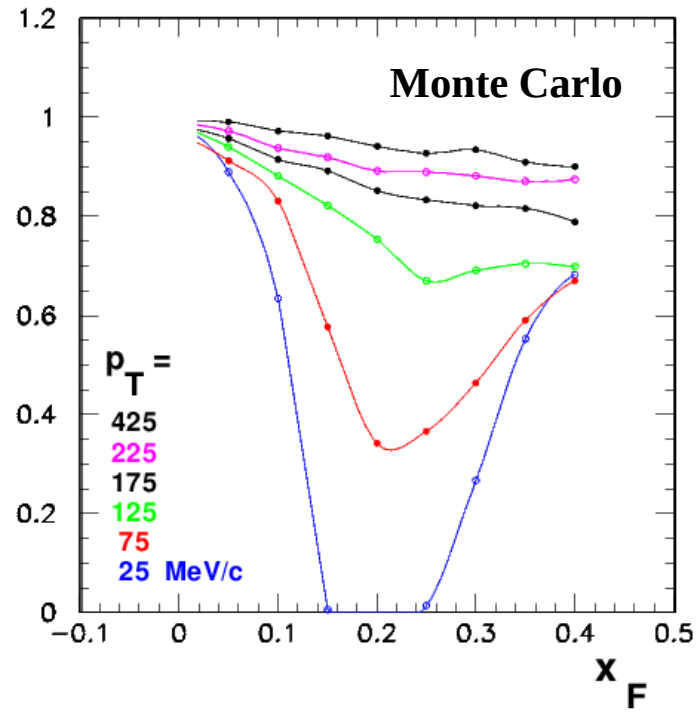
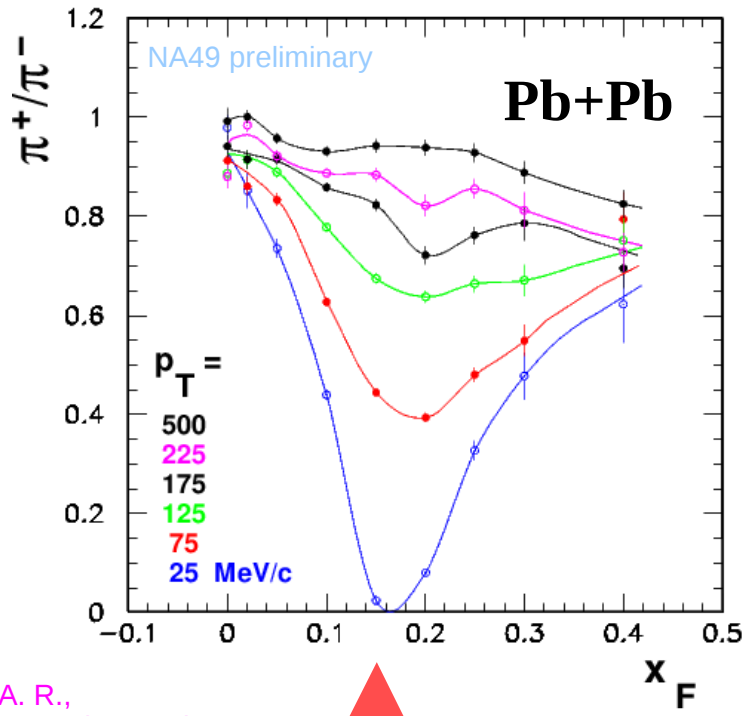
$\sqrt{s}_{NN} = 17.3 \text{ GeV}$



by I. Sputowska

- Charged spectators in non-central collisions generate **electromagnetic fields**.
- Can we use them as a new source of information on the space-time evolution of the system ?
- Can we use resonances ?

## ***2) EM effects in heavy ion collisions***



**NA49,**  
 $\sqrt{s}_{NN} = 17.3$  GeV  
**Pb+Pb, peripheral**

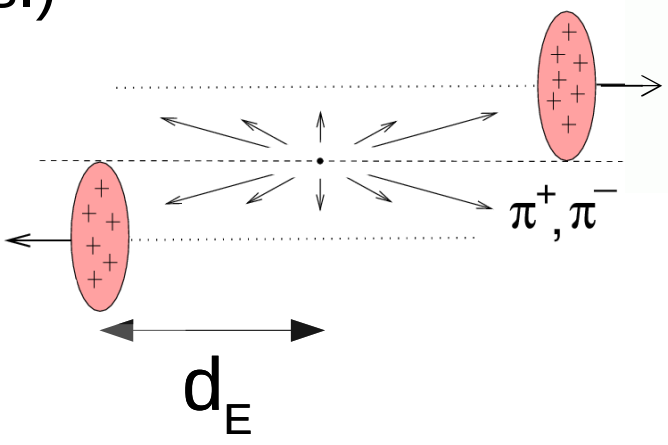
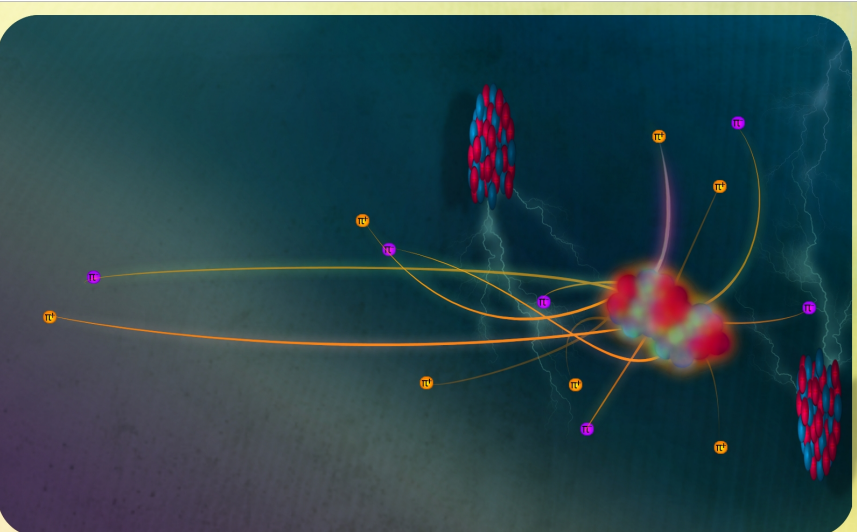
A. R.,  
 Acta Phys. Polon.  
 B42 (2011) 867

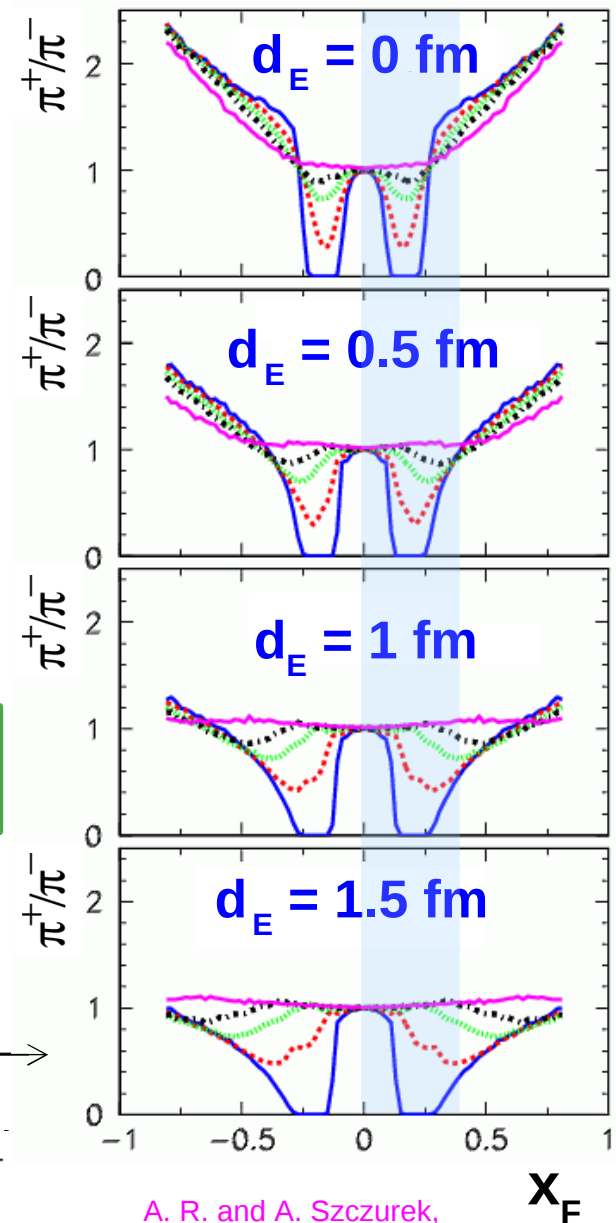
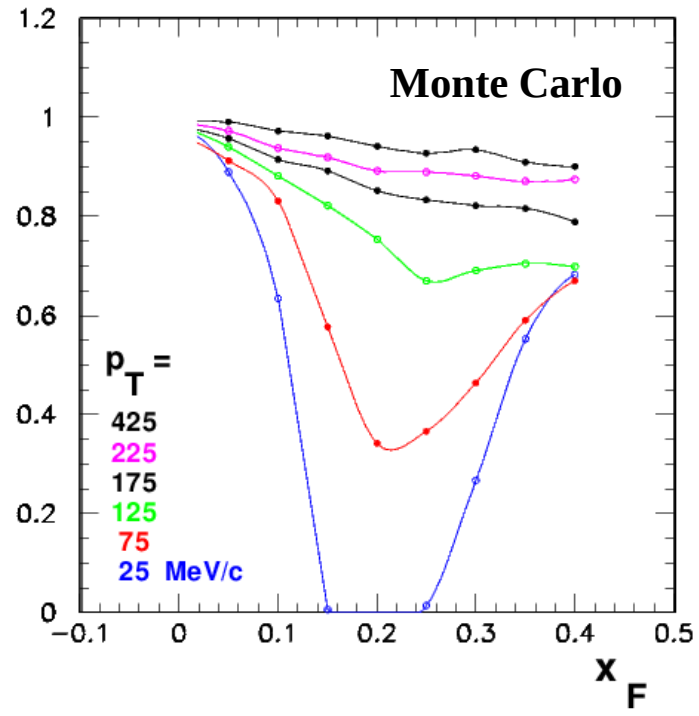
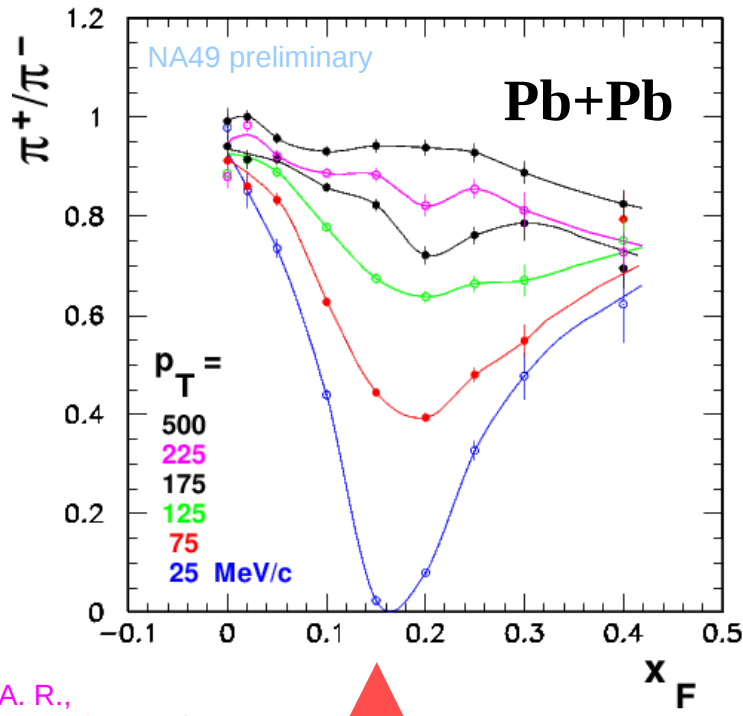
↑  
**spectator  
 velocity**

$$x_F = \frac{p_L}{p_L^{beam}}$$

(c.m.s.)

**$d_E \approx 0.75$  fm !**





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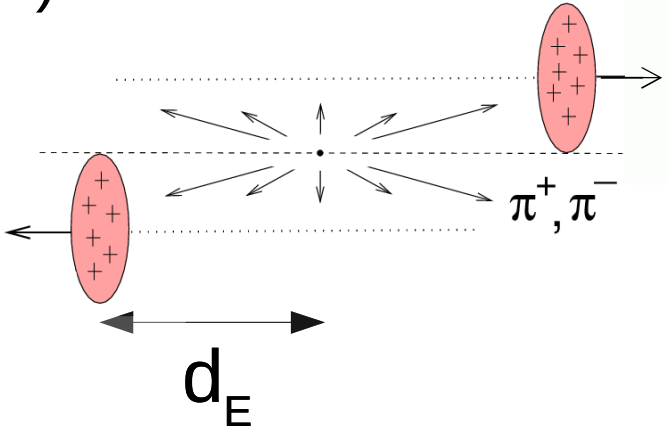
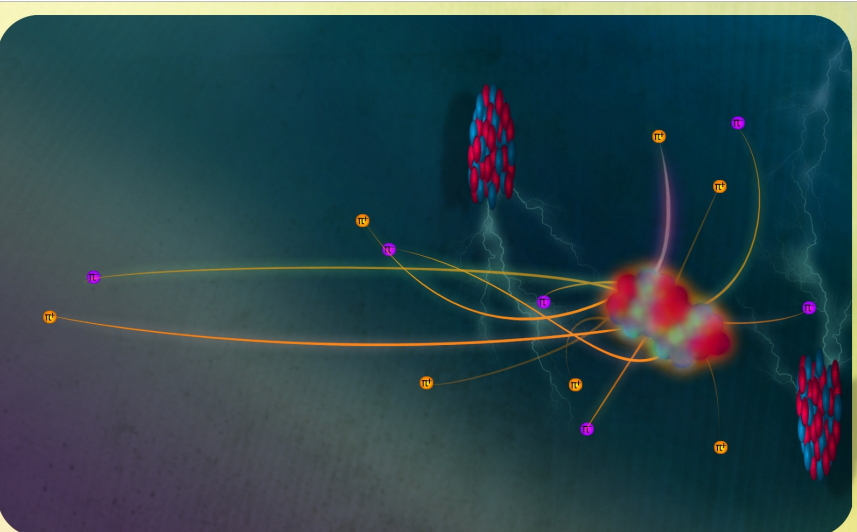
↑  
**spectator  
velocity**

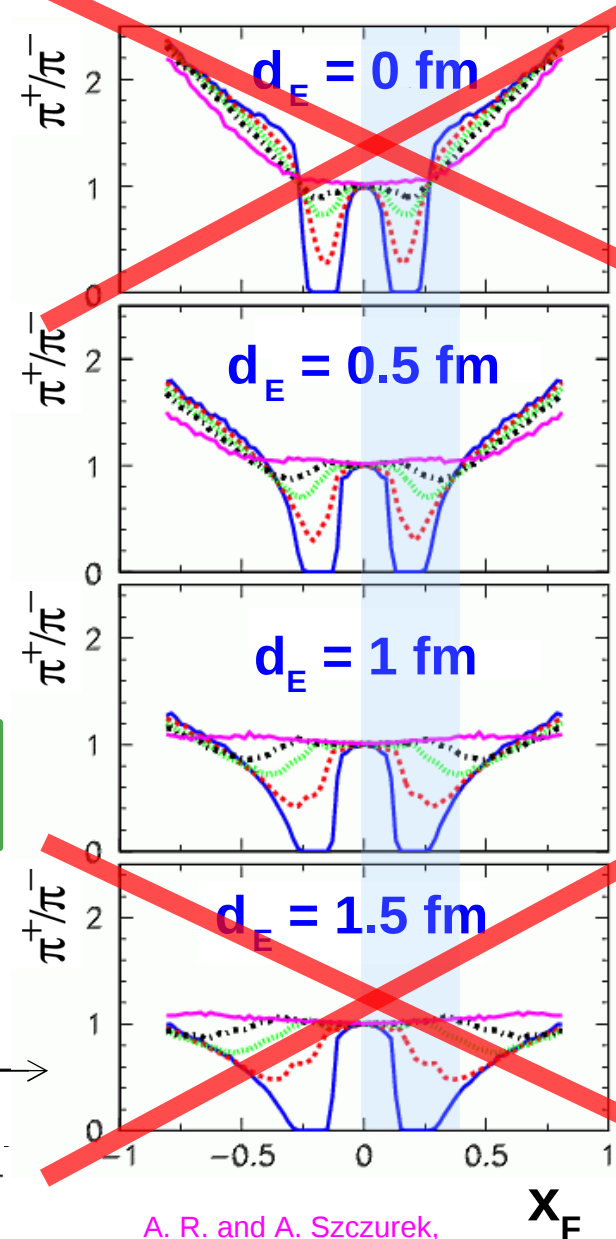
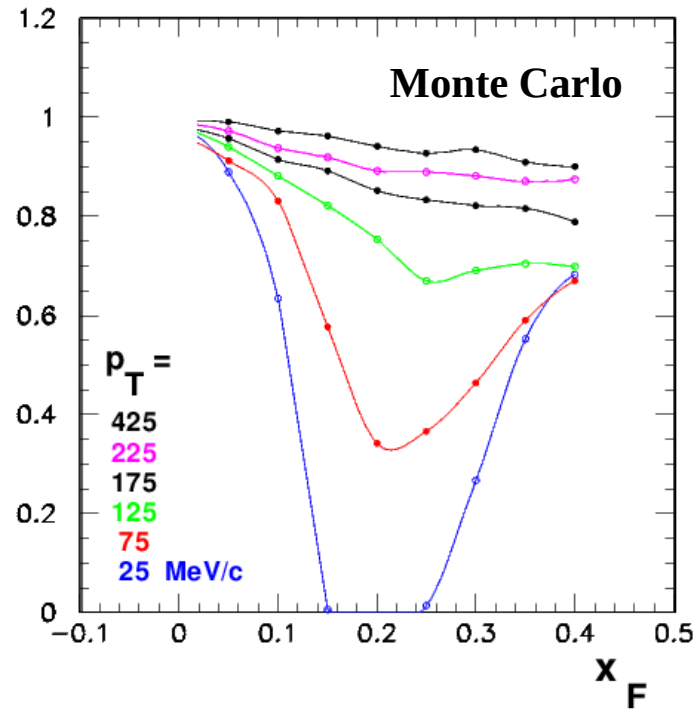
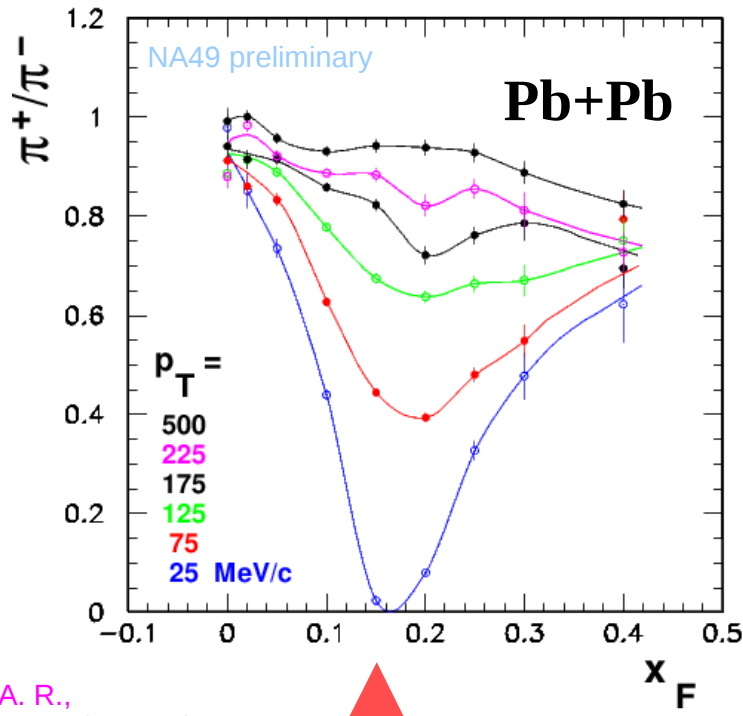
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A. R. and A. Szczurek,  
Phys. Rev. C75 (2007)  
054903





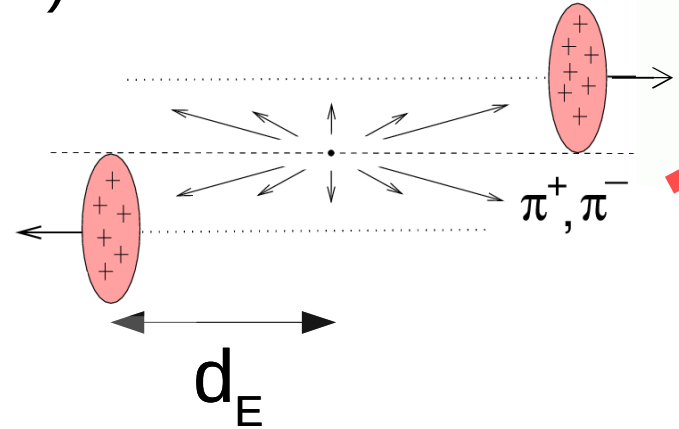
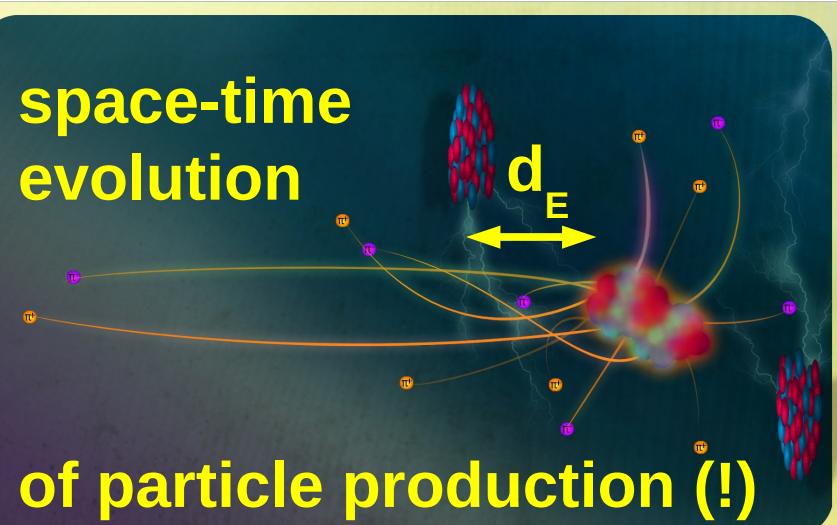
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↑  
**spectator  
velocity**

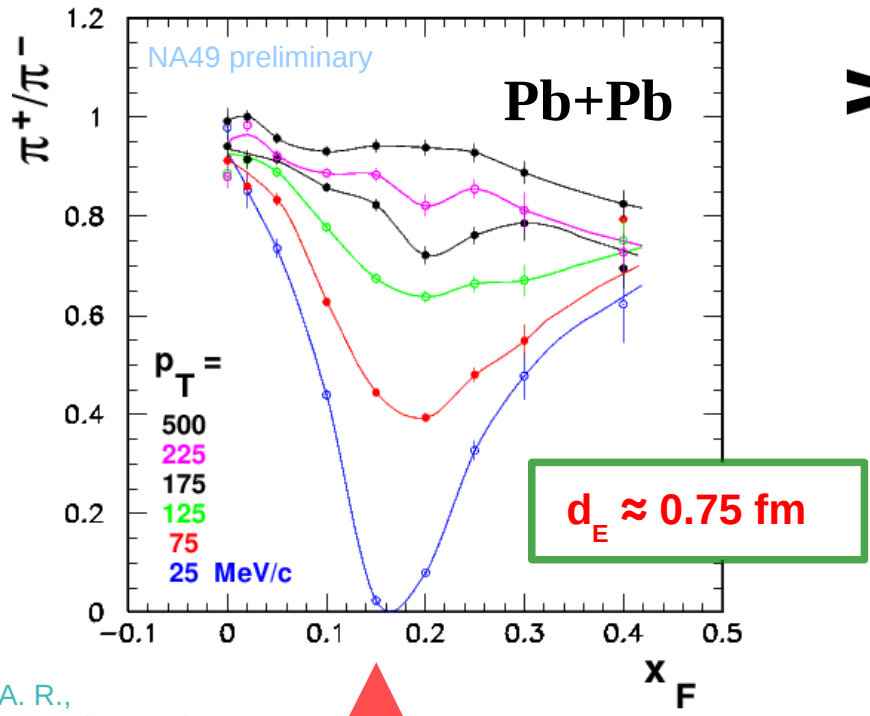
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(c.m.s.)

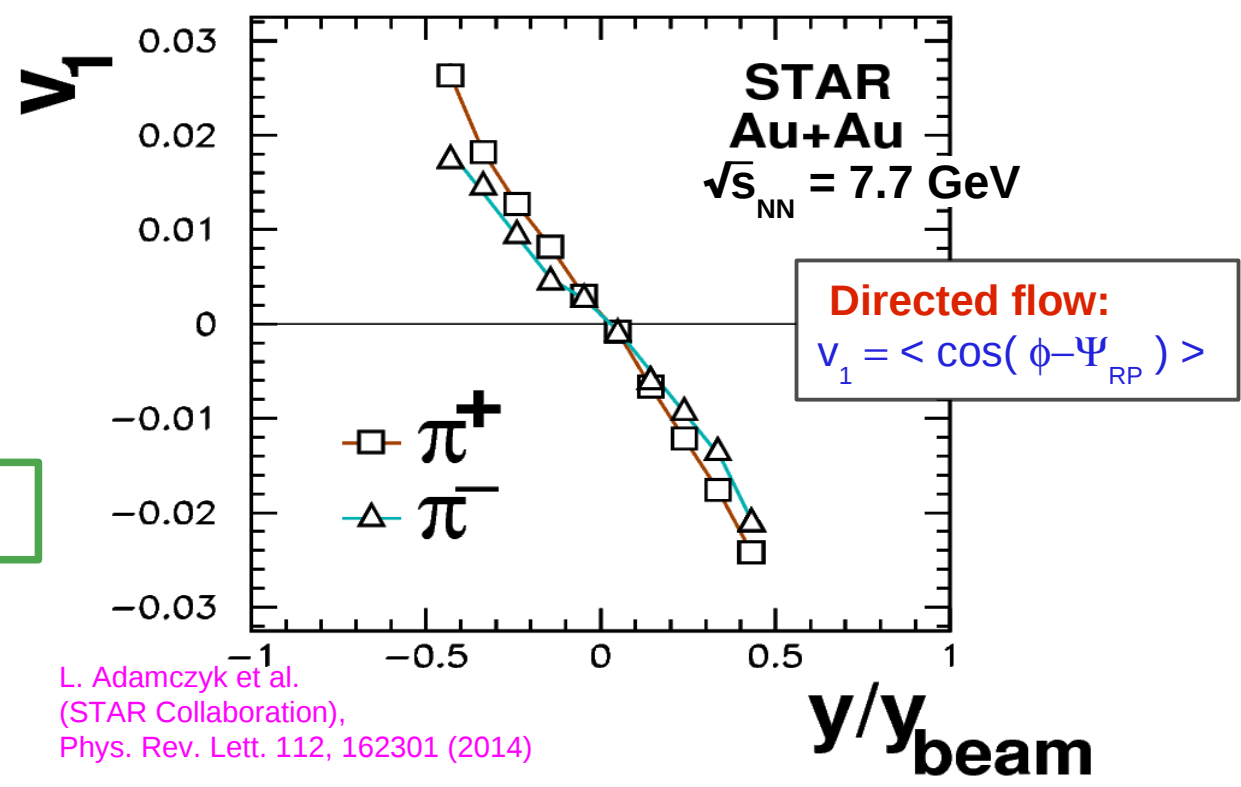
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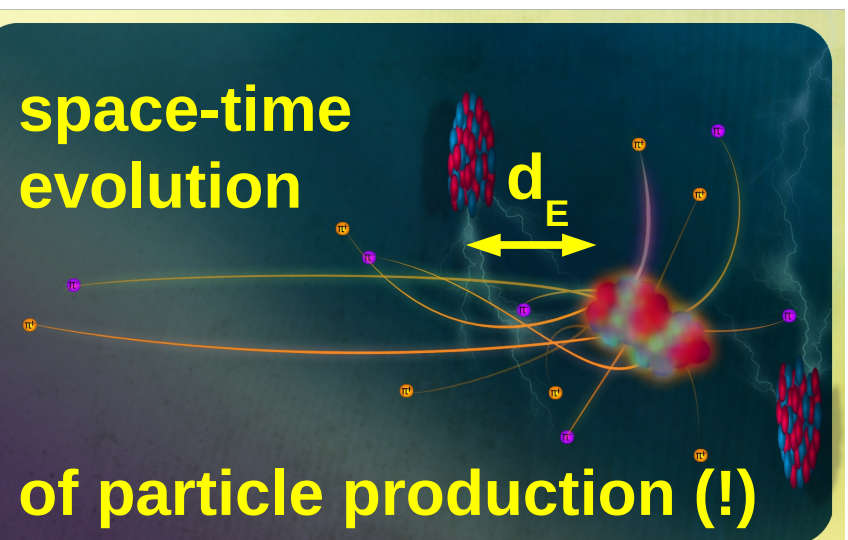
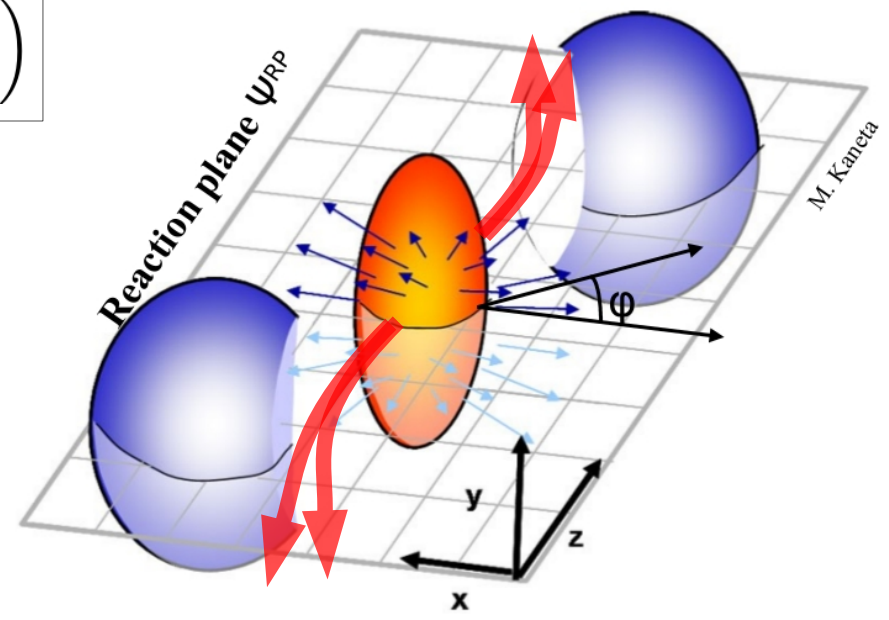


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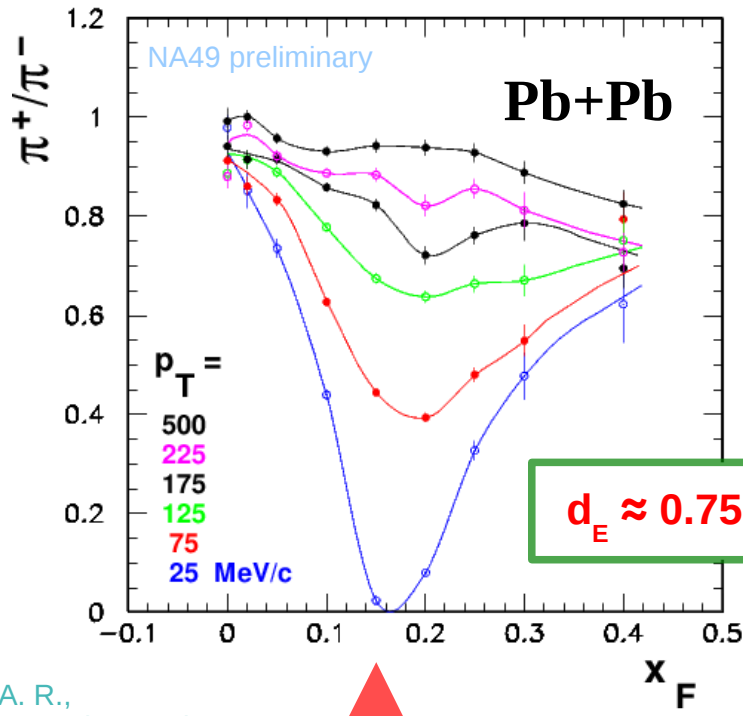
L. Adamczyk et al.  
(STAR Collaboration),  
Phys. Rev. Lett. 112, 162301 (2014)

$$y = \frac{1}{2} \ln \left( \frac{E+p_L}{E-p_L} \right)$$

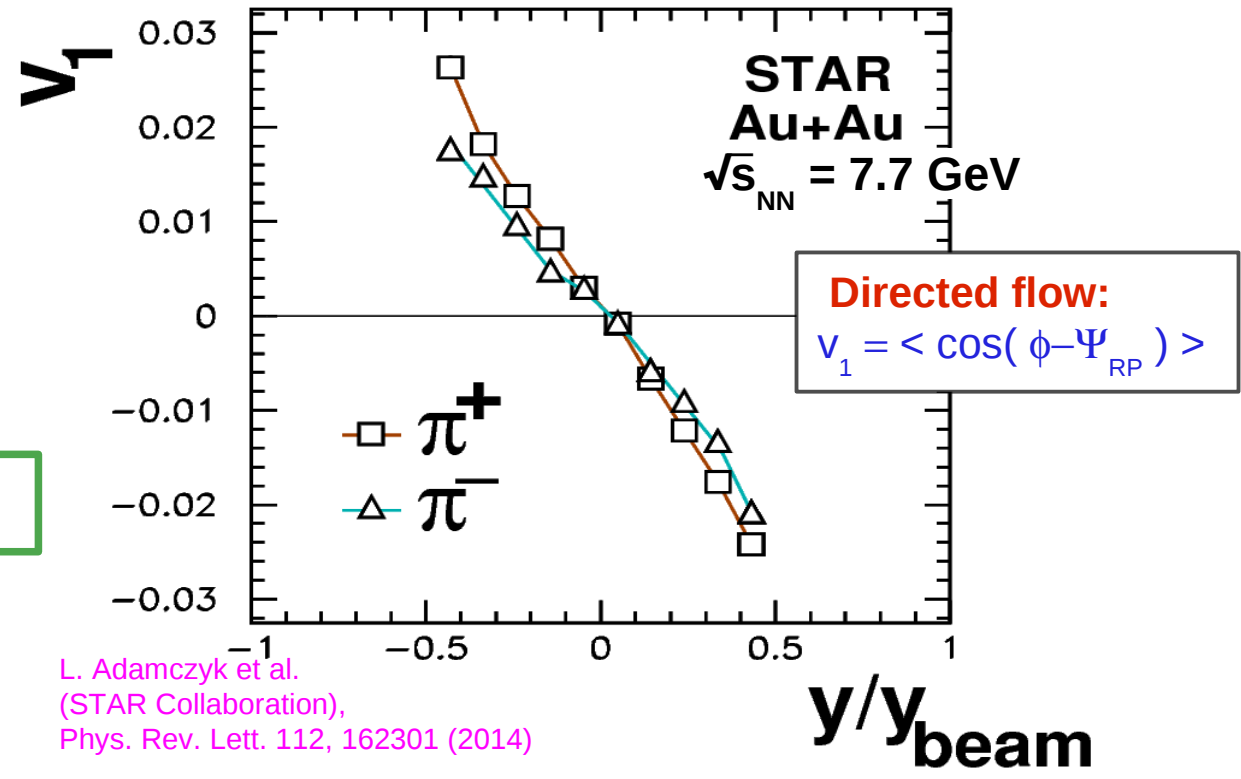


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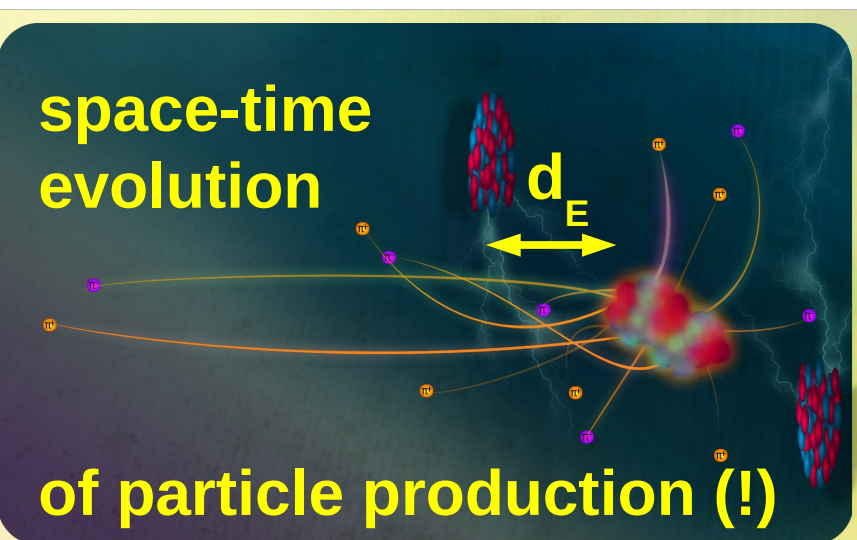
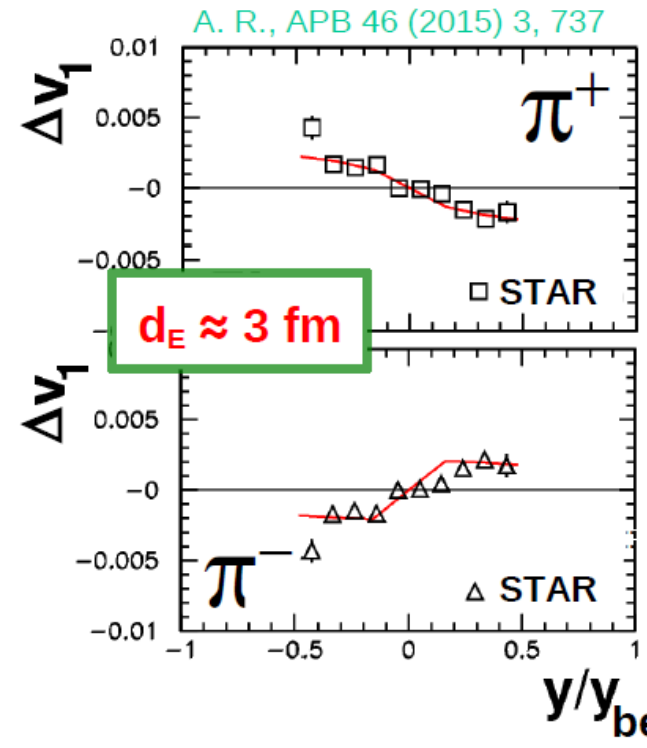




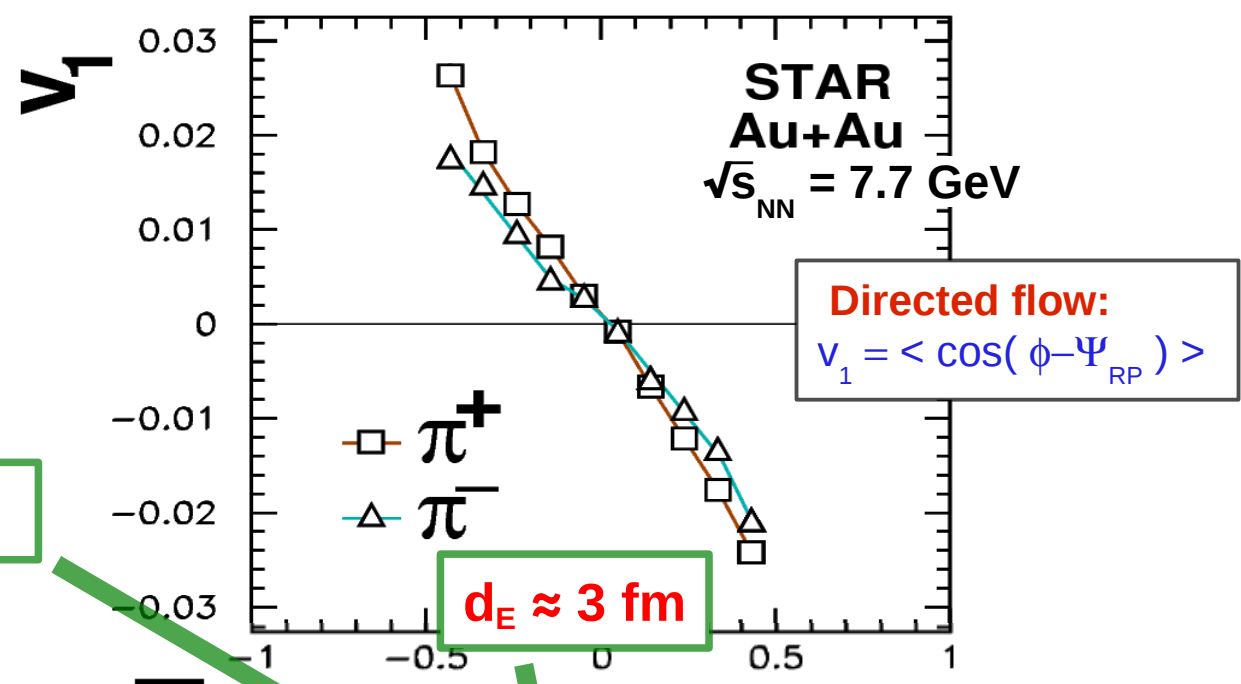
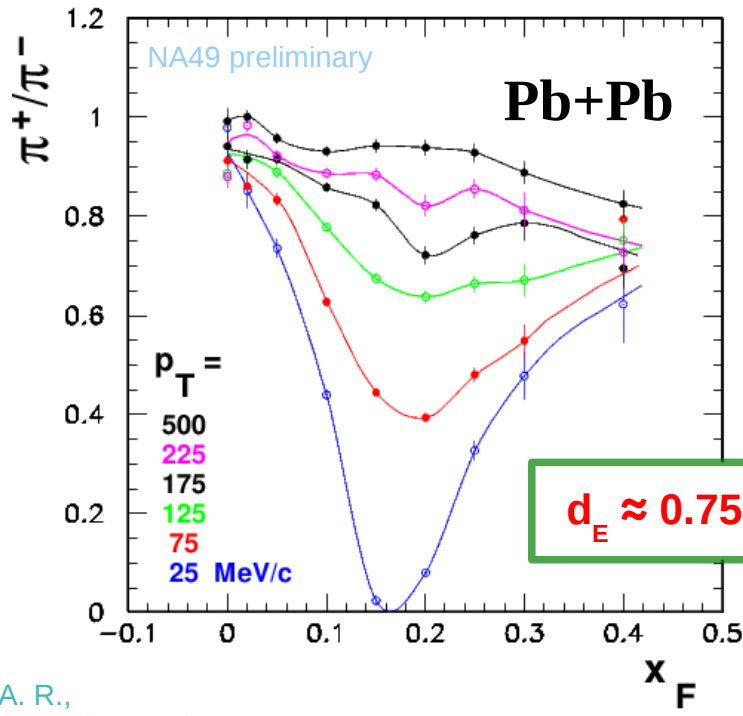
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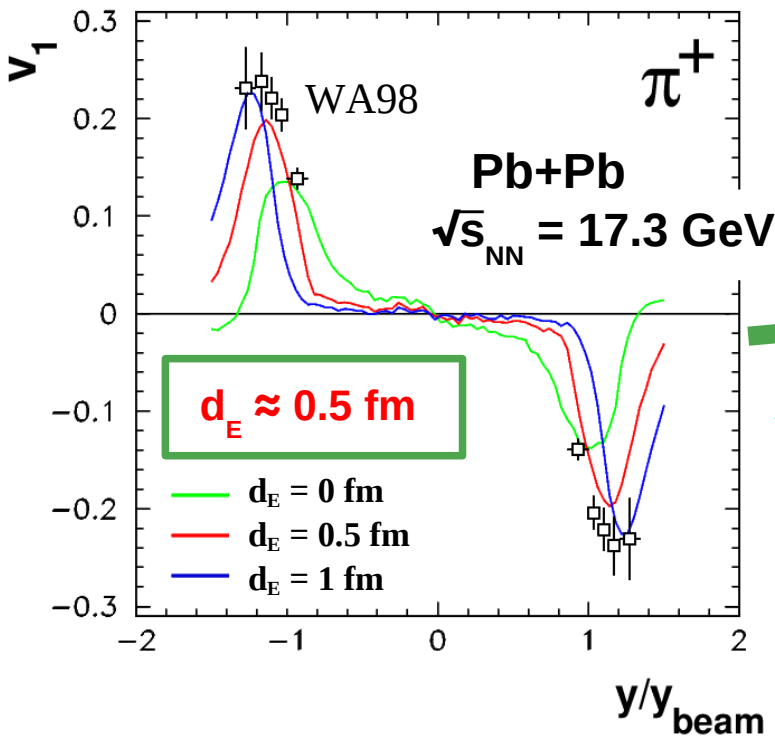
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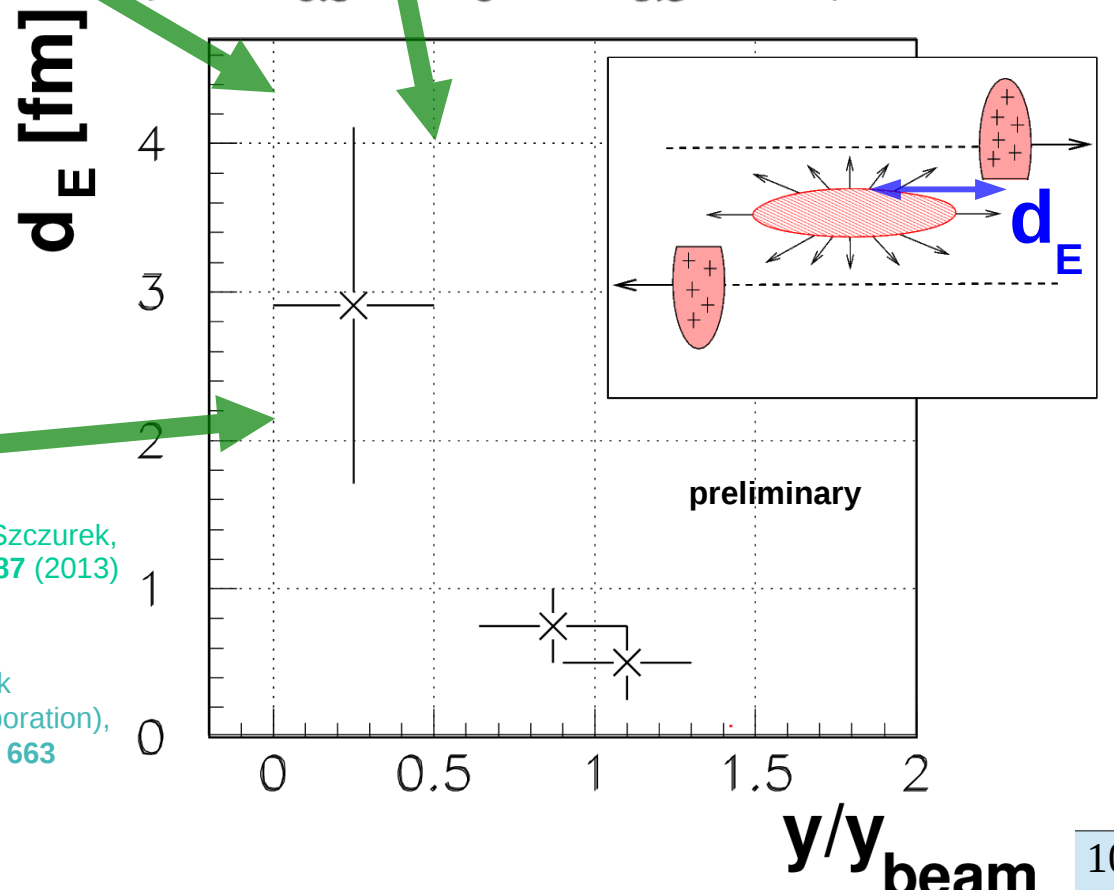


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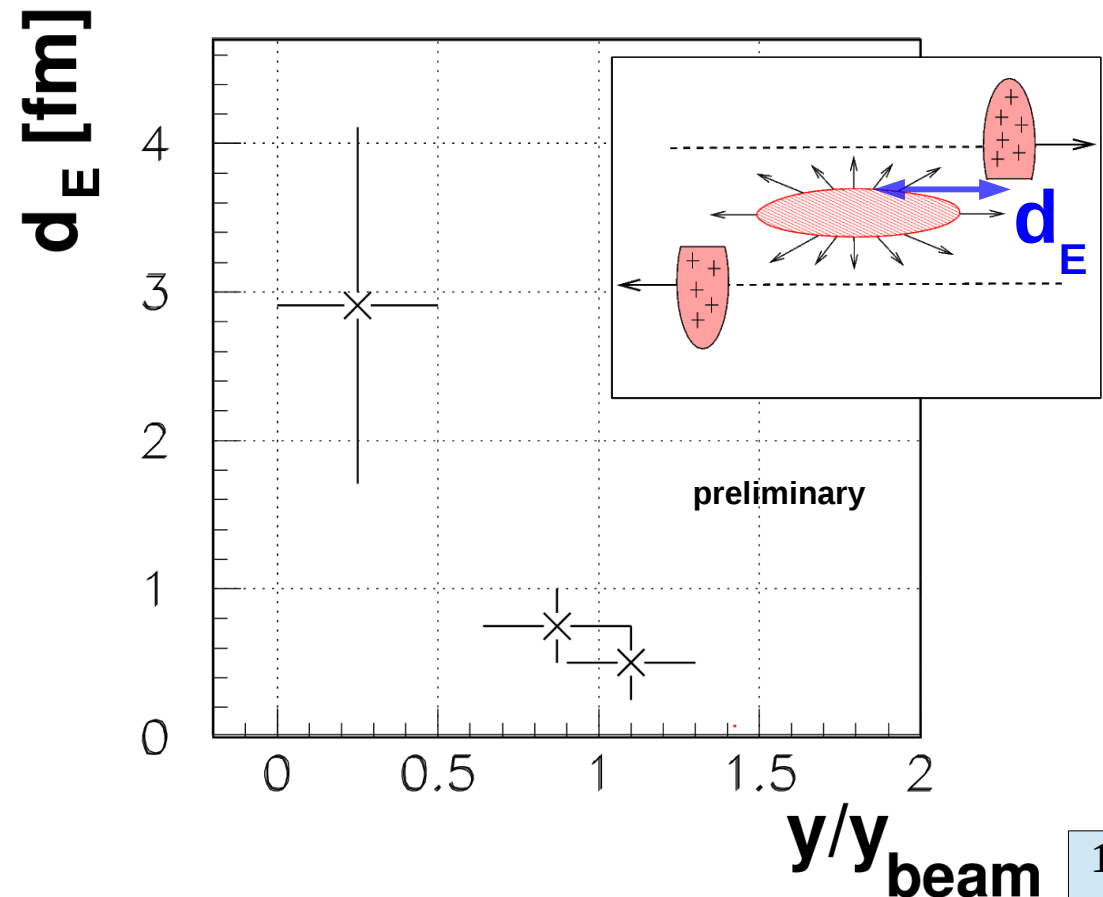


A. R. and A. Szczurek,  
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054909.

H. Schlagheck  
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Nucl. Phys. A 663  
(2000) 725.



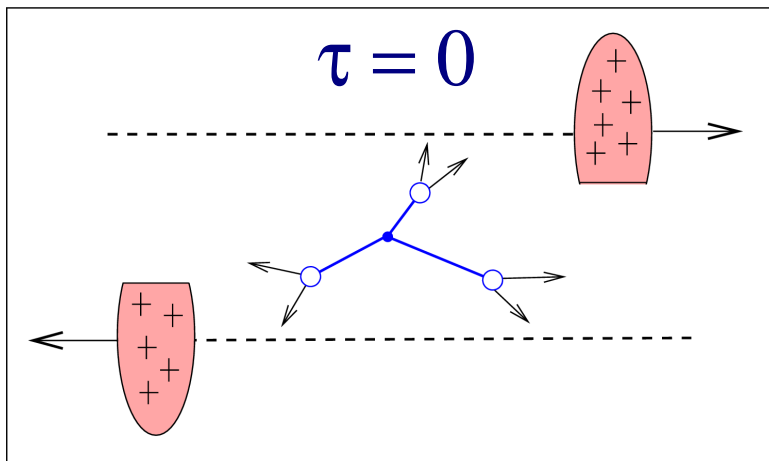
### 3) Space-time evolution of the system ...



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## Toy Monte Carlo model:

- pion production from resonances ;
- $\Delta \rightarrow \rho\pi$  and  $\rho \rightarrow \pi\pi$  ;
- $(y, p_T)$  spectra  $\sim$  known in p+p ;
- baryon stopping of the  $\Delta$  ;
- Breit-Wigner's, lifetimes, etc.



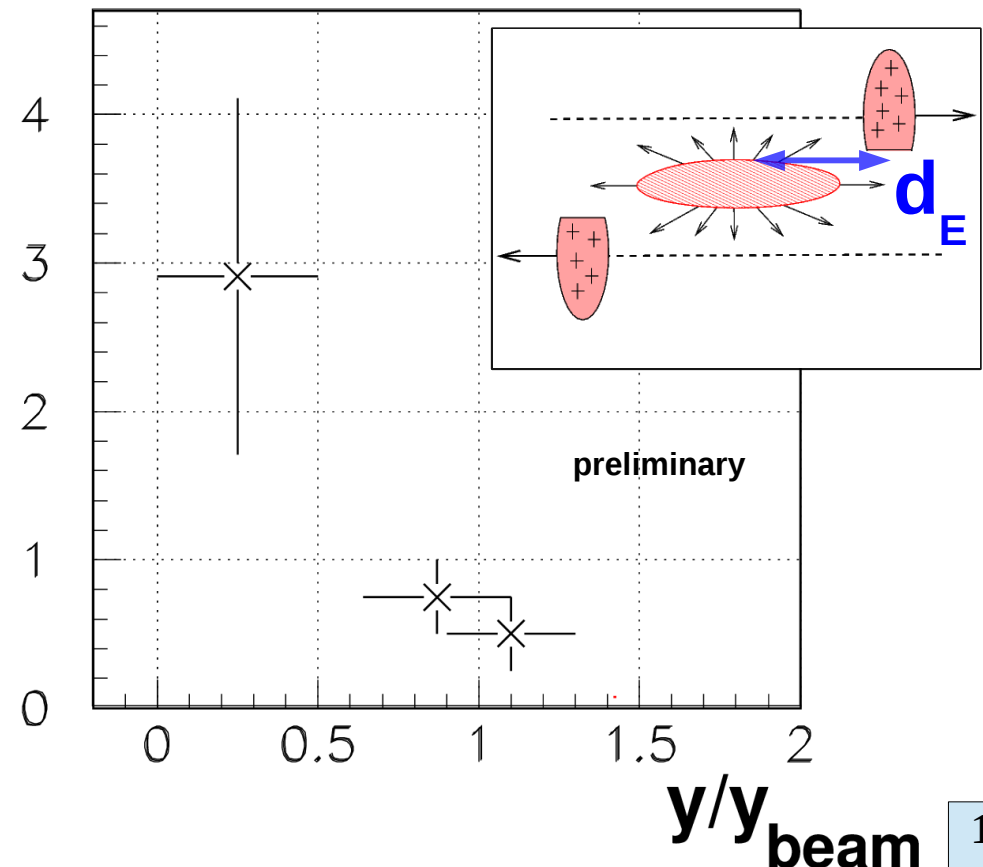
many thanks to

**K. Redlich** **M. Rózańska**  
**L. Leśniak** **H.G. Fischer**

Input:

- K.Fiałkowski, W.Kittel, Rept. Prog. Phys 46 (1983) 1283.
- T.Anticic et al., Phys. Rev. C86 (2012) 054903.
- M.Aguilar-Benitez et al., Z. Phys. C 50 (1991) 405.
- D. Drijard et al., Z. Phys. C 21 (1984) 321.
- D.E.Groom et al., Eur. Phys. C. 15 (2000) 1.
- A.R., CERN-THESIS-2003-005, and references therein.

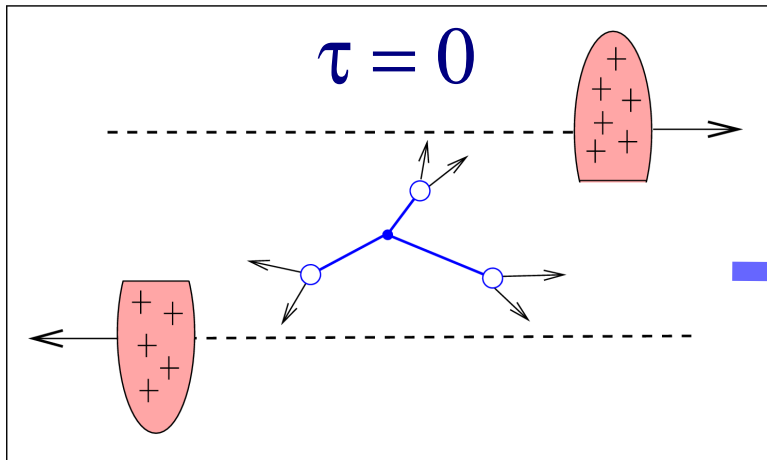
$d_E$  [fm]



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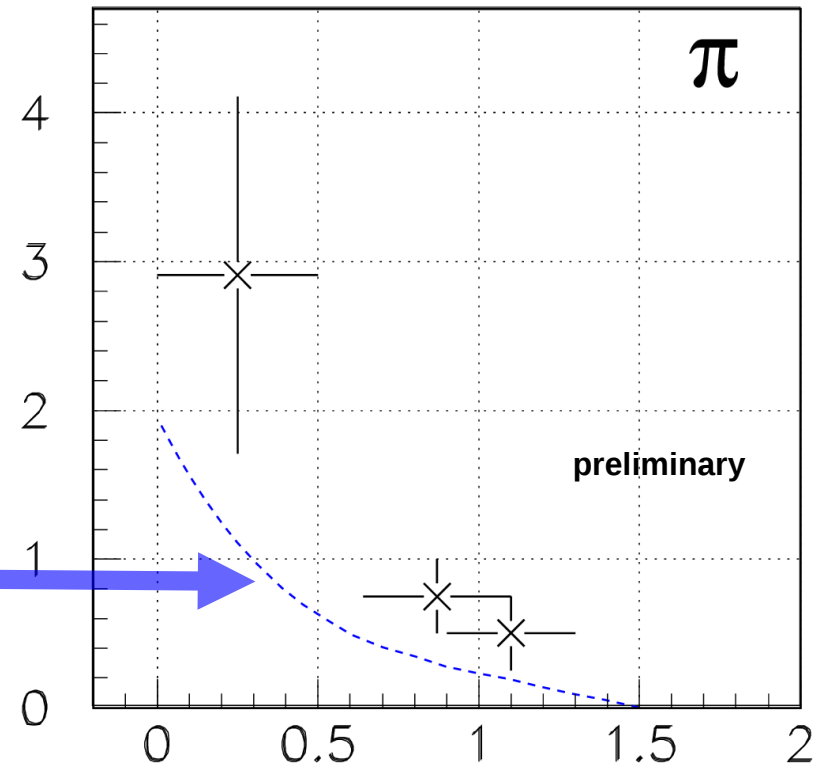
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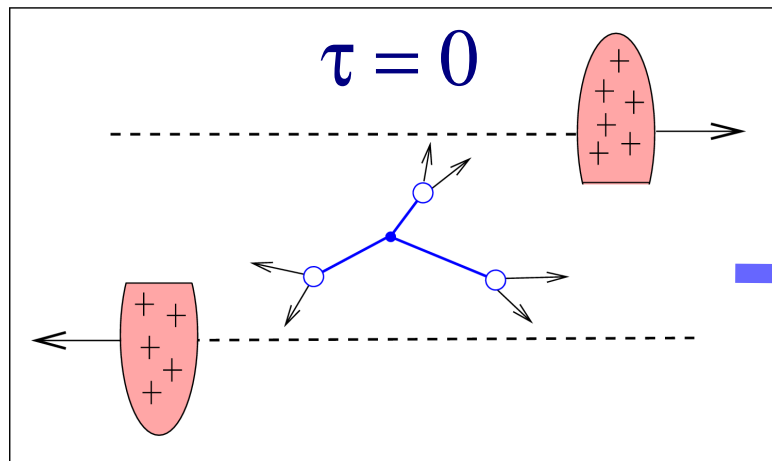
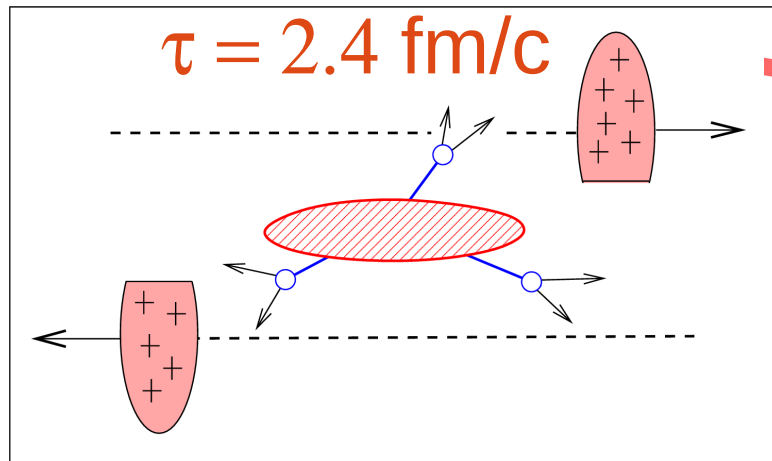


$\pi$

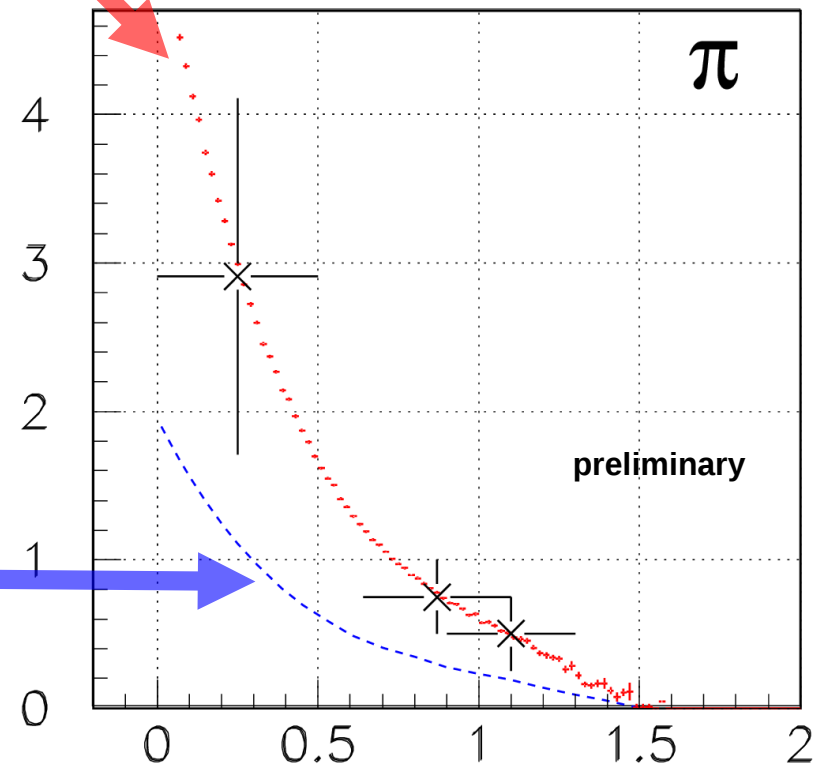
preliminary

$y/y_{\text{beam}}$

### 3) Space-time evolution of the system ...

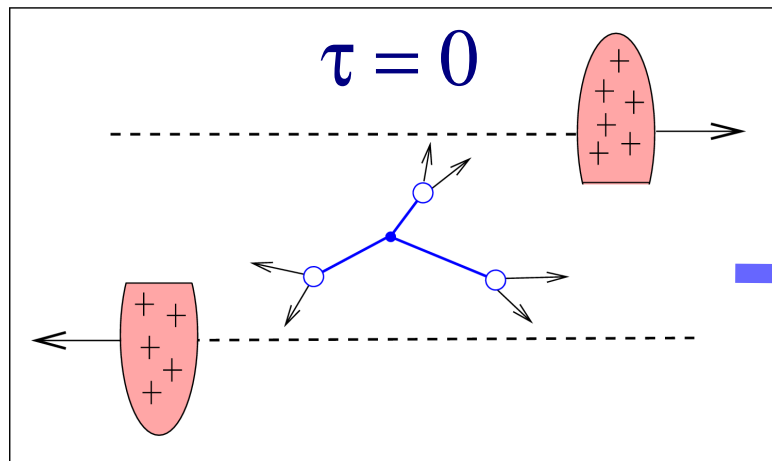
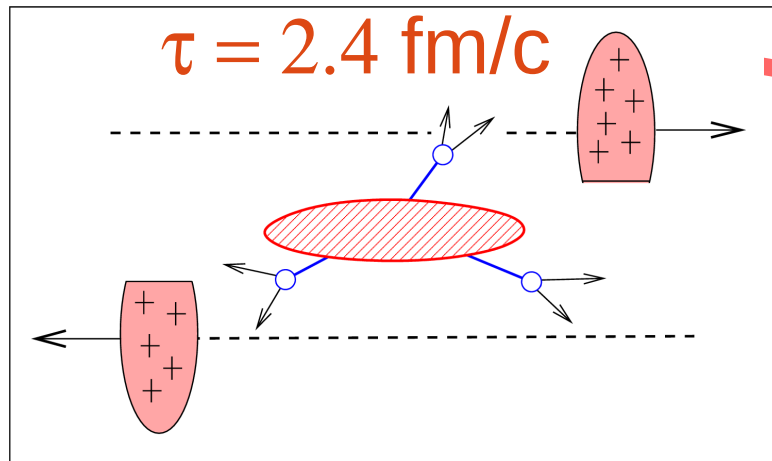


$d_E$  [fm]

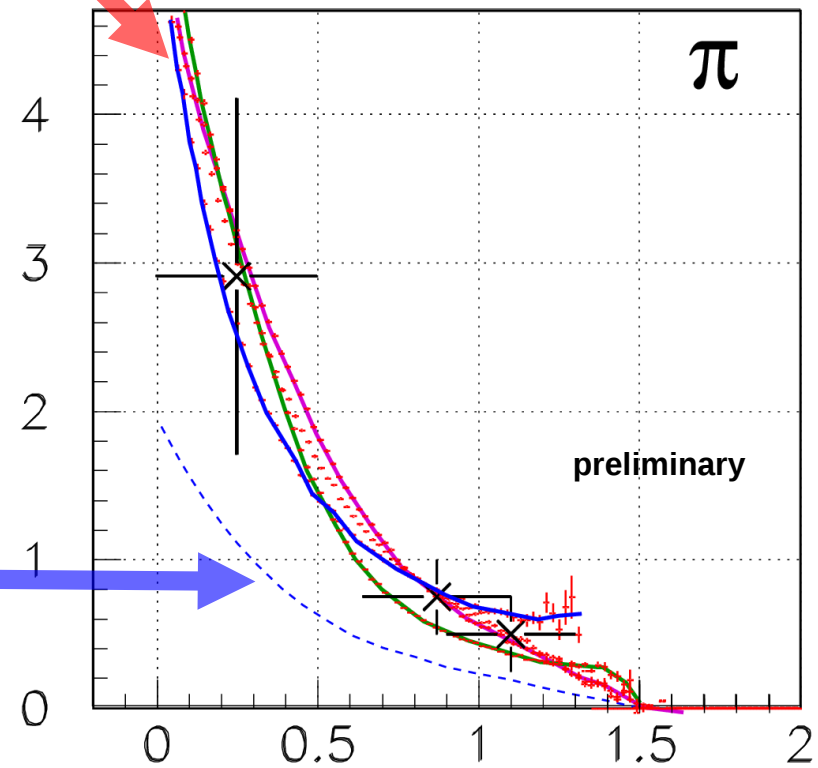


$y/y_{\text{beam}}$

### 3) Space-time evolution of the system ...

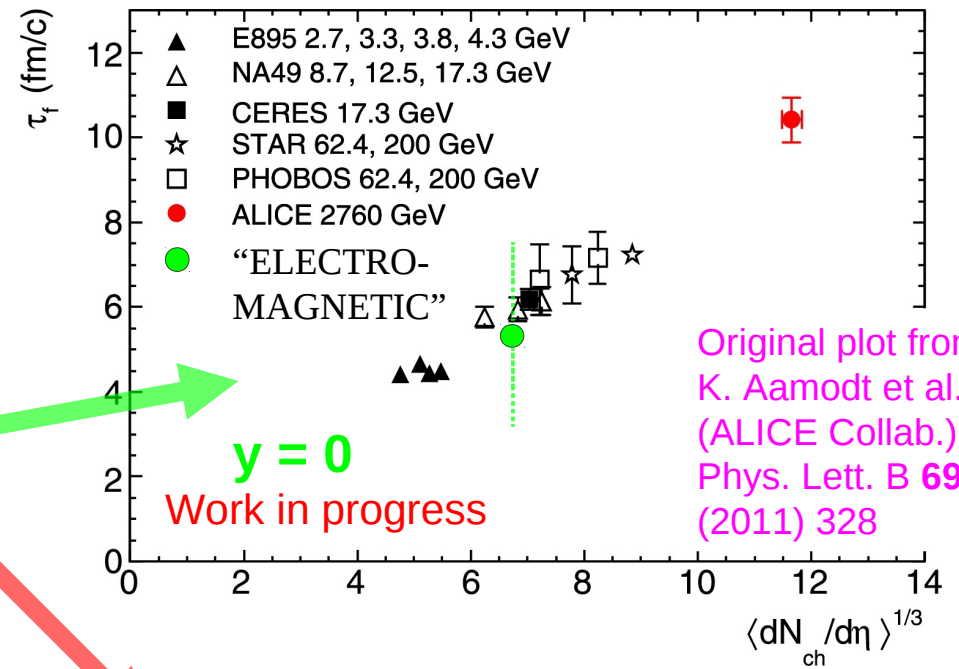
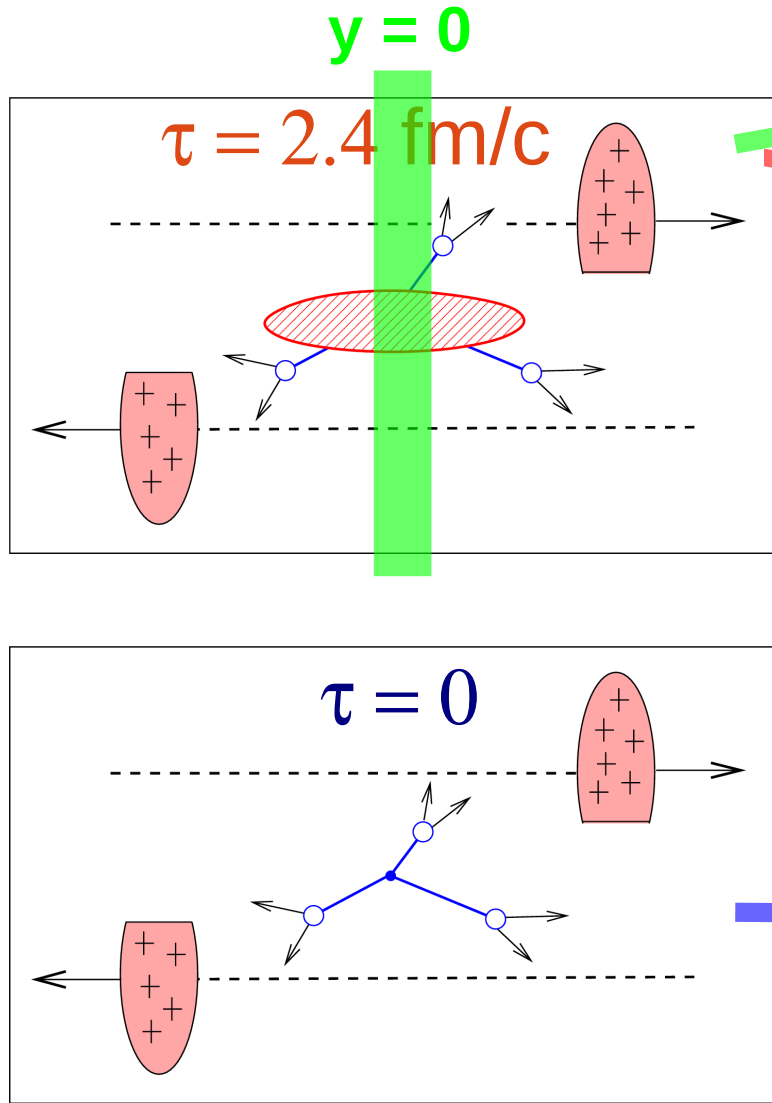


$d_E$  [fm]

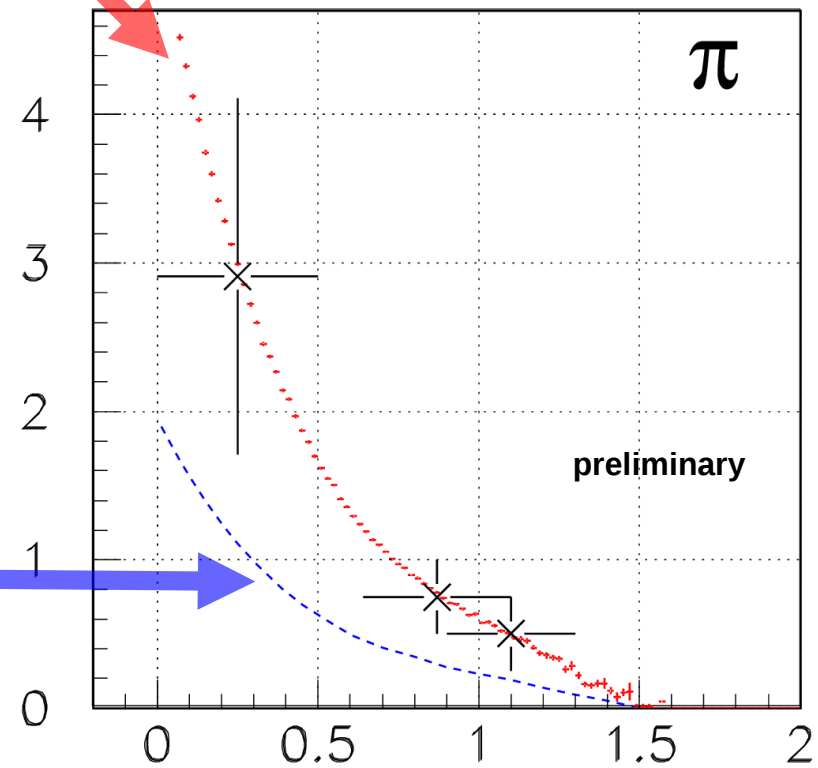


$y/y_{\text{beam}}$

# 3) Space-time evolution



$d_E \text{ [fm]}$



$y/y_{\text{beam}}$



## 4) Summary & outlook

EM effects in heavy ion collisions are sensitive to the distance  $d_E$  between the pion emission site and the spectator(s).

They can be used as a new source of information on the longitudinal space-time evolution of the system.

-----  
Plan (2015-2020):

/ built up a group, joined  
NA61/SHINE @ CERN /

1. to get more data on these effects  
(NA61/SHINE, SPSC-P-330-ADD-8, NICA, EPJA vol. 52 (2016)) ;
2. to clarify the situation at LHC.

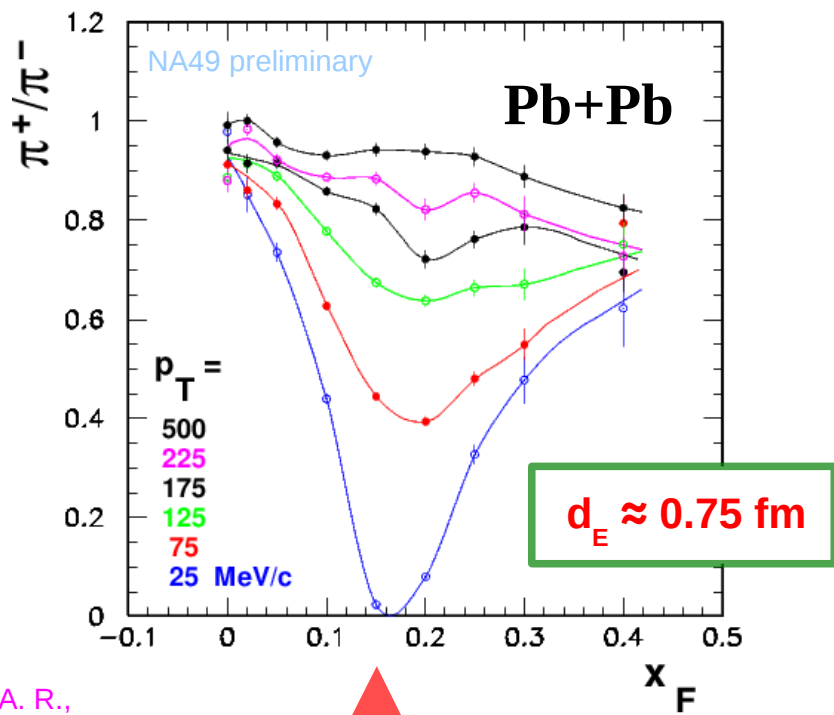
Help, advice and discussion are more than welcome.

Thank you !

# Acknowledgments.

This work was supported by the National Science Centre, Poland (grant no. 2014/14/E/ST2/00018).

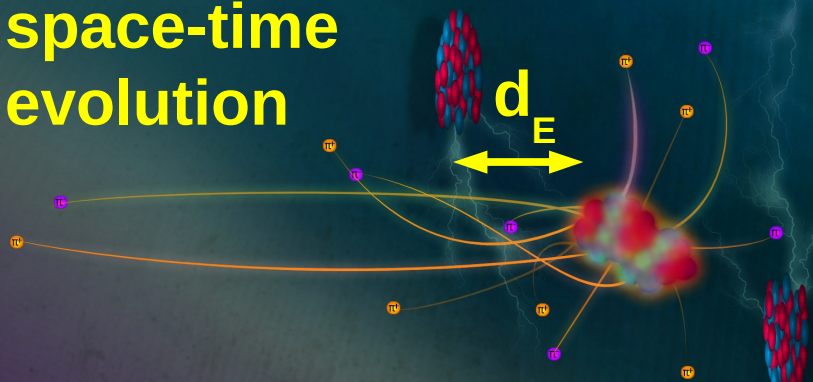
# *Extra slides*



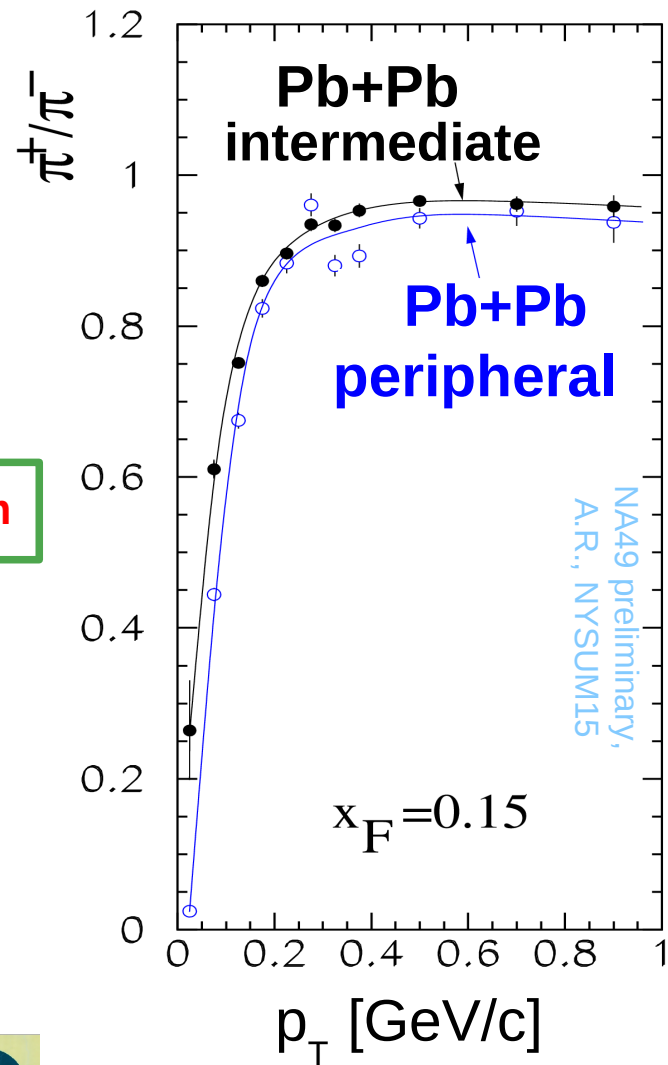
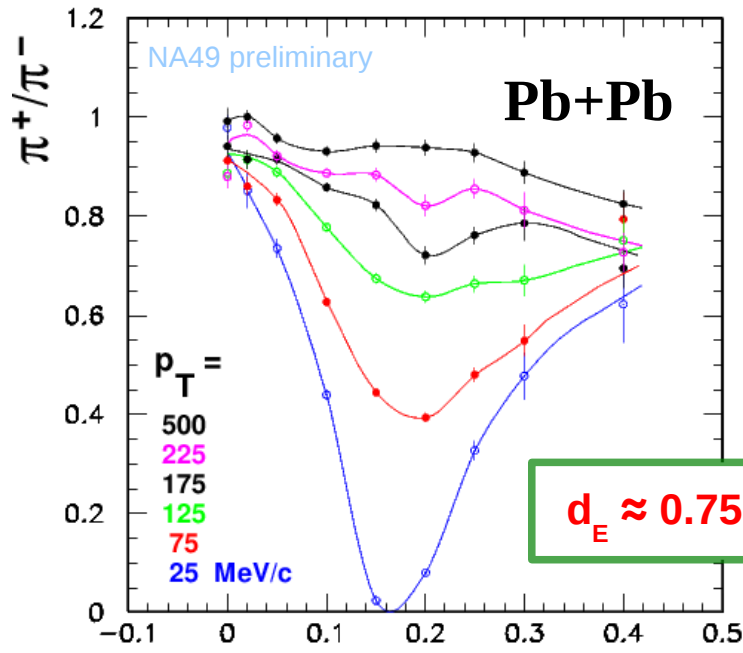
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$y = y_{\text{beam}}$

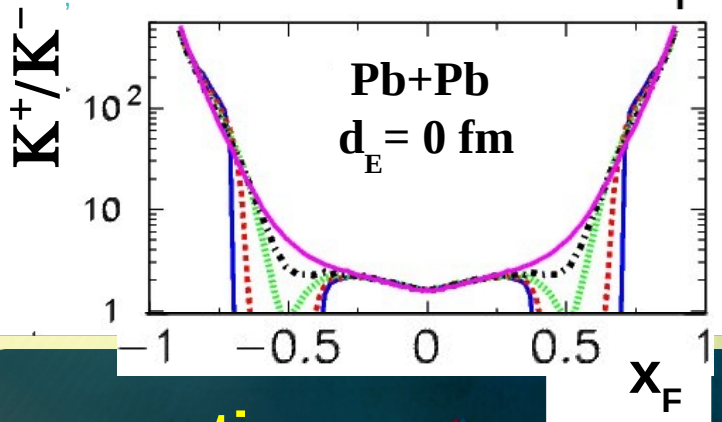
space-time  
 evolution



of particle production (!)

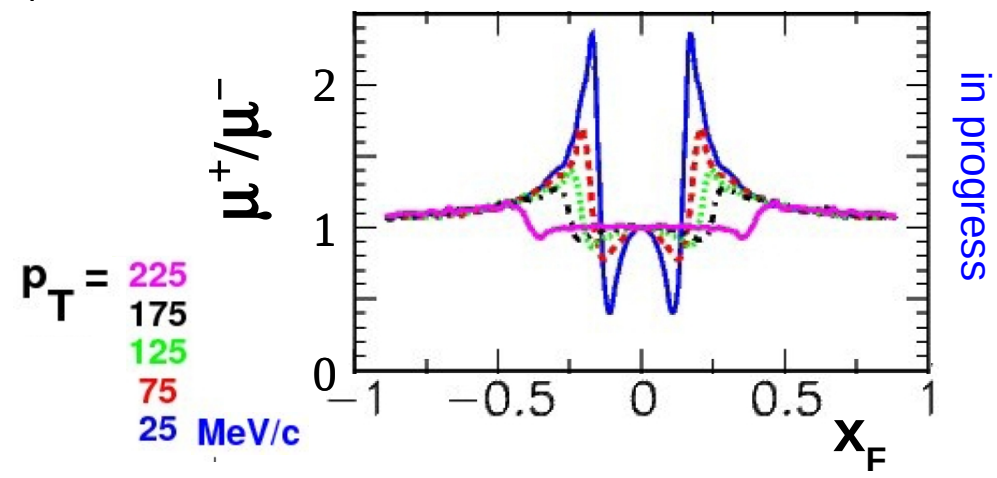
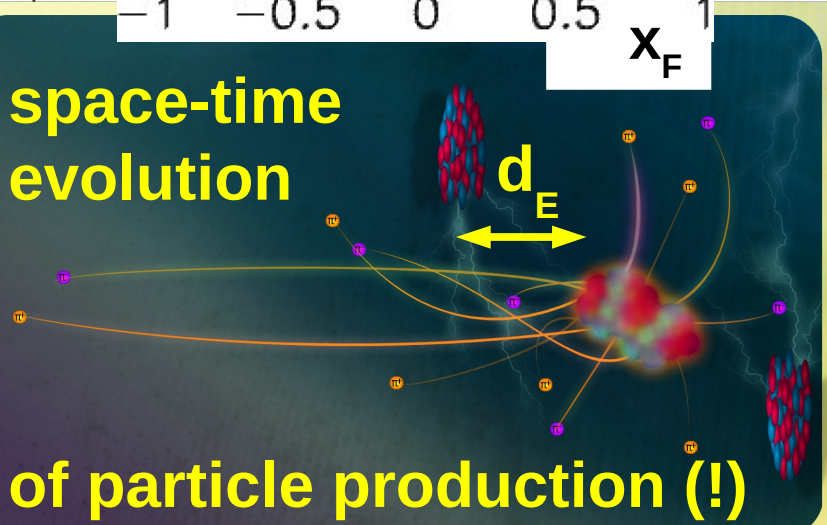


**Pb+Pb,  
SPS,  
17.3 GeV**

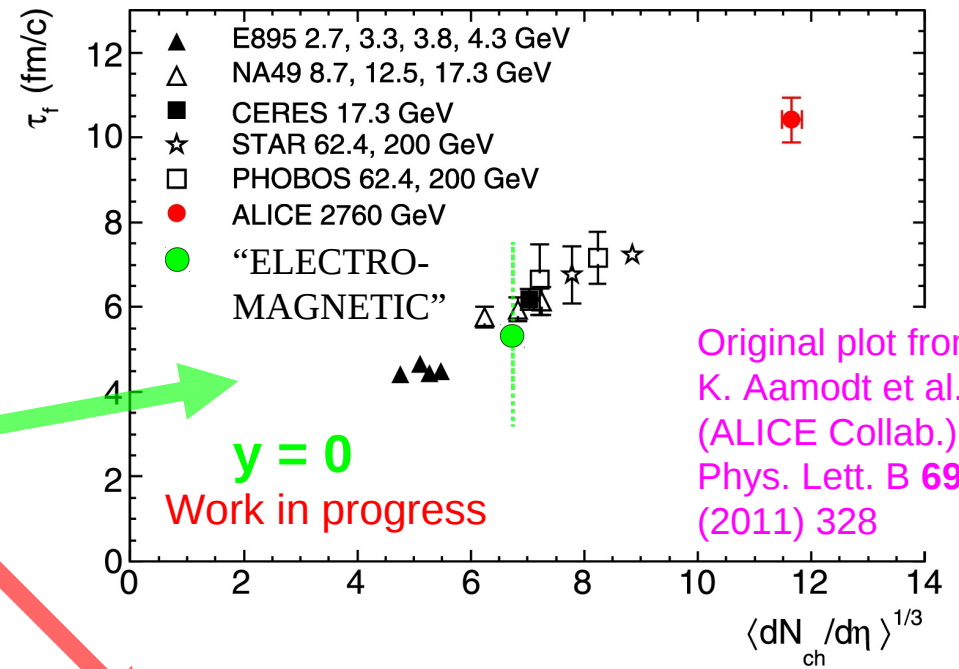
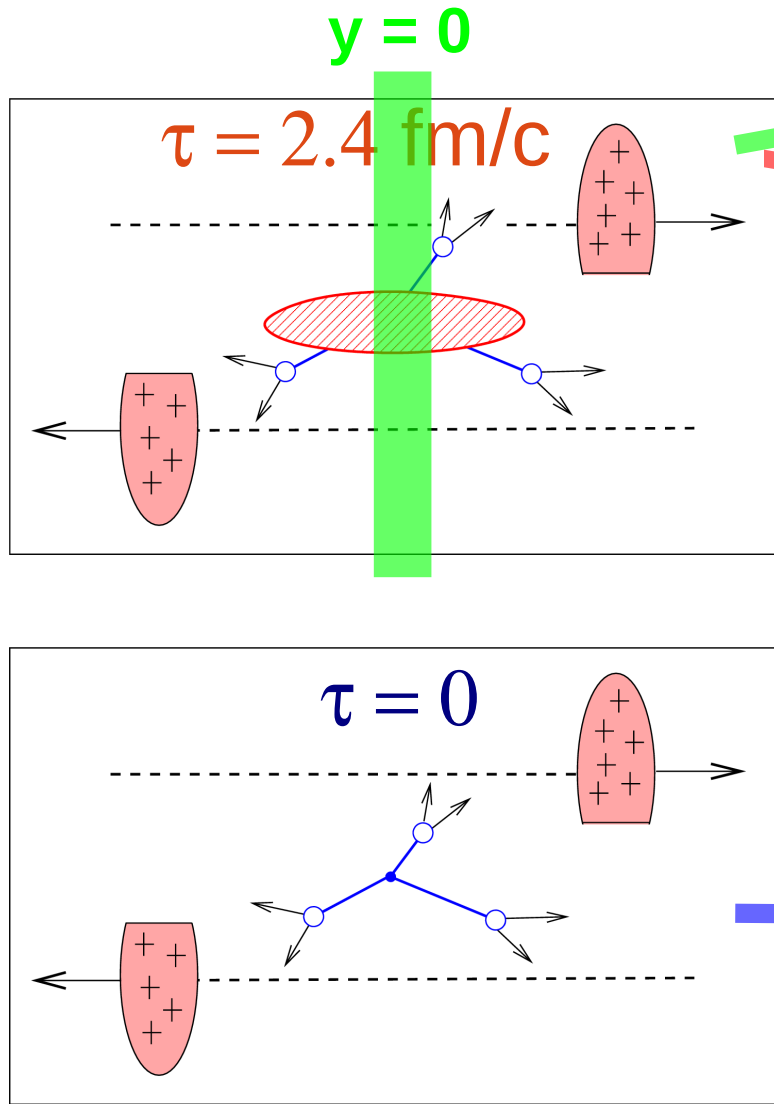


**UPC, Au+Au,  
200 GeV  
( $\gamma\gamma \rightarrow \mu^+\mu^-$ )**

**b=112 fm,  $d_E = 0$  fm**

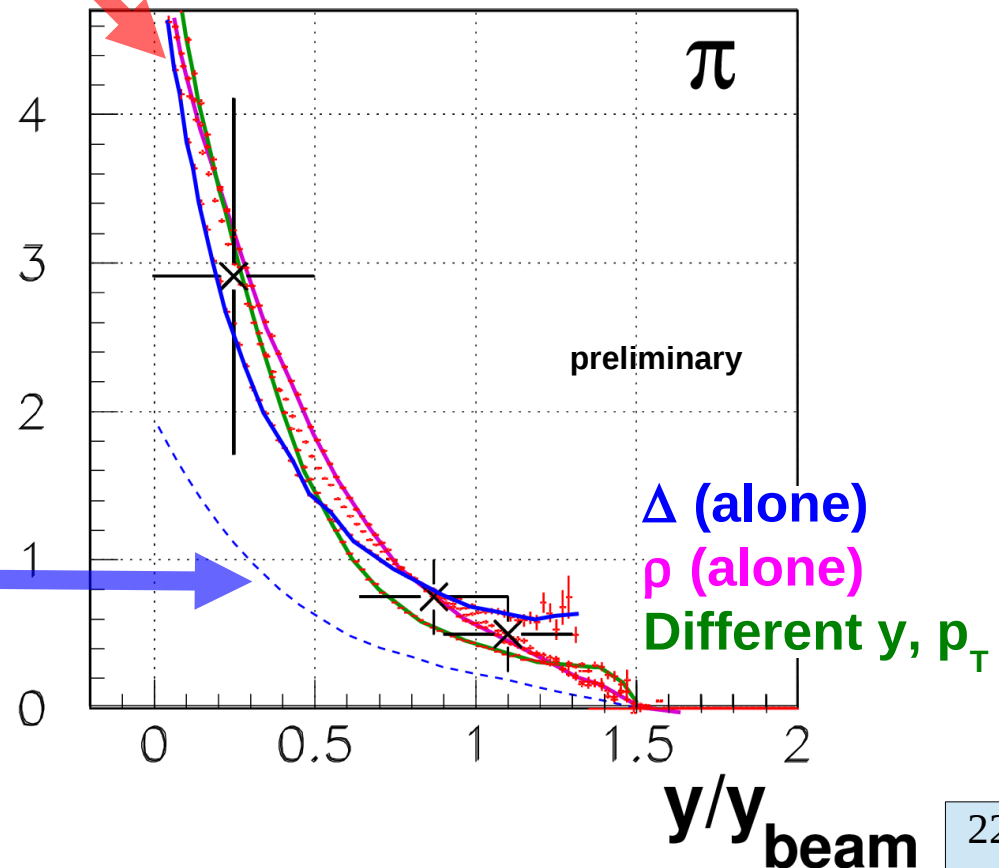


# 3) Space-time evolution

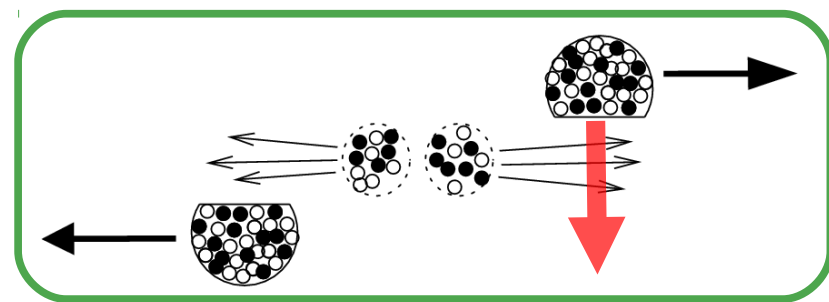


Original plot from:  
 K. Aamodt et al.  
 (ALICE Collab.),  
 Phys. Lett. B **696**  
 (2011) 328

**$d_E \text{ [fm]}$**



# NA49, 158 A GeV/c

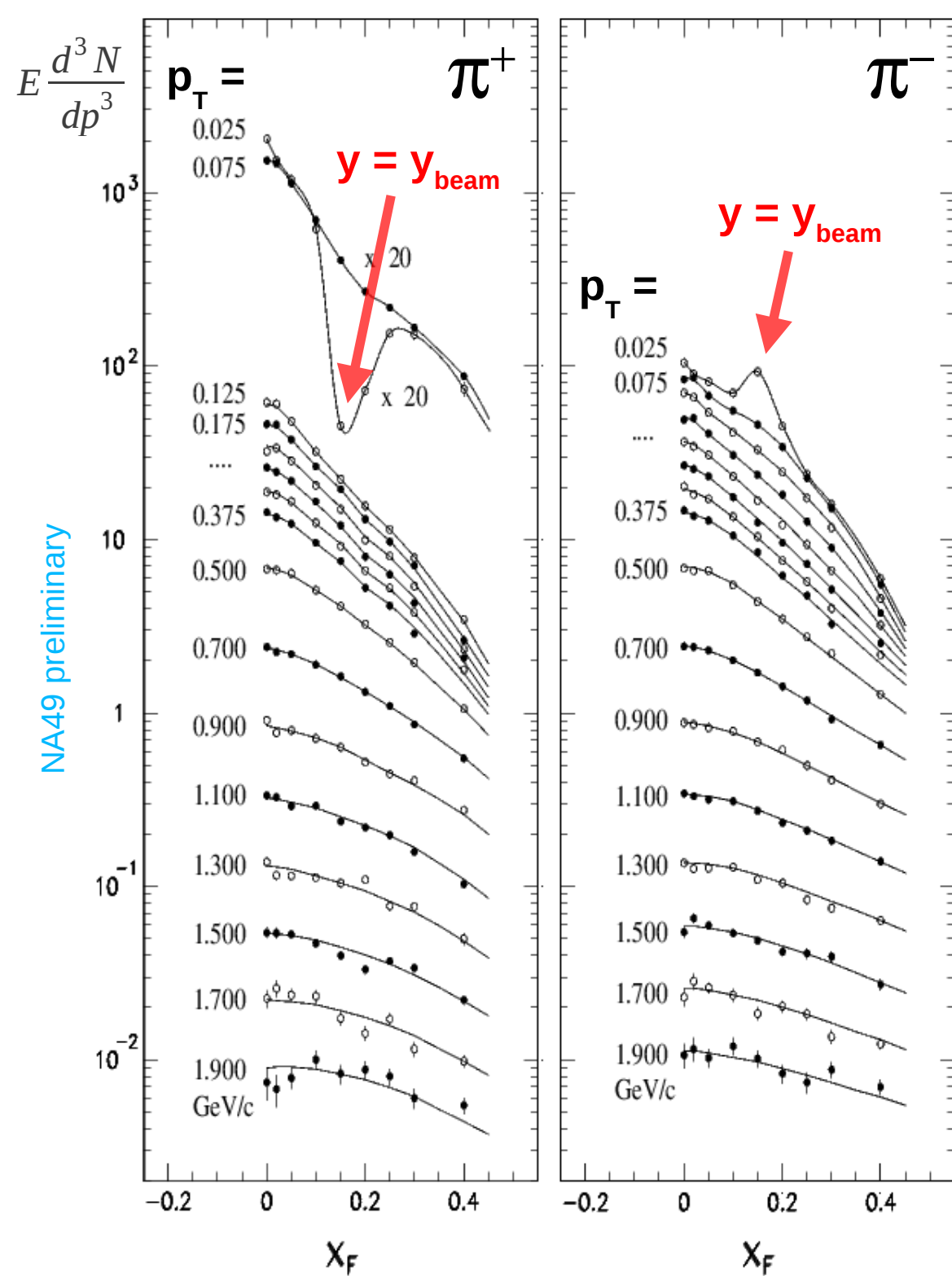


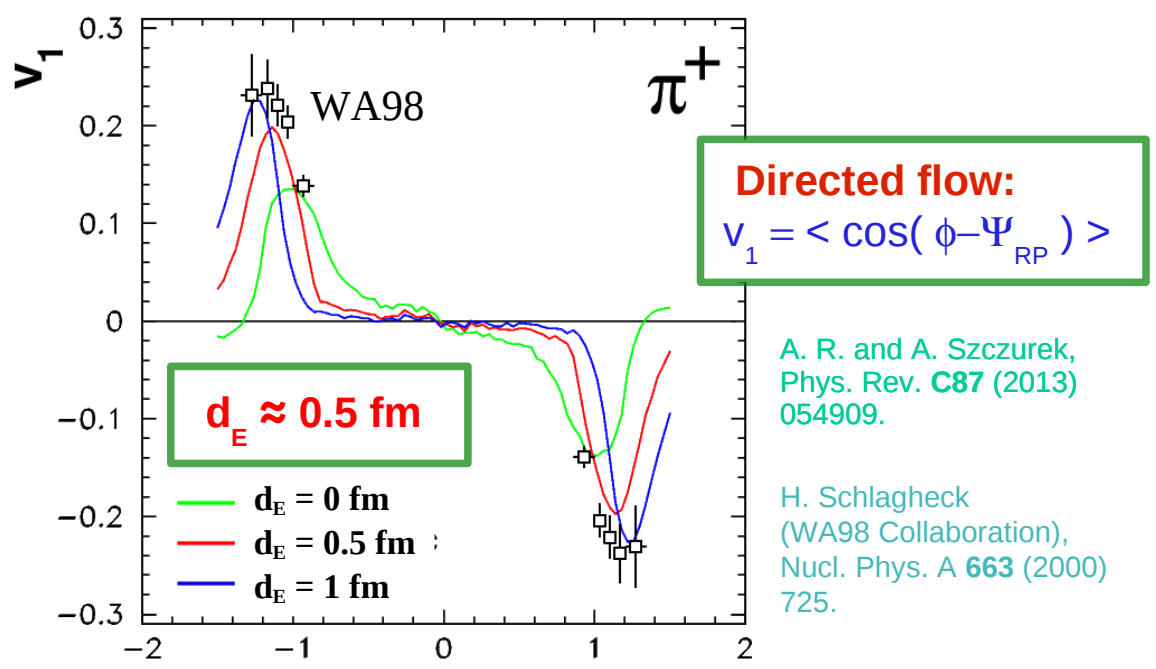
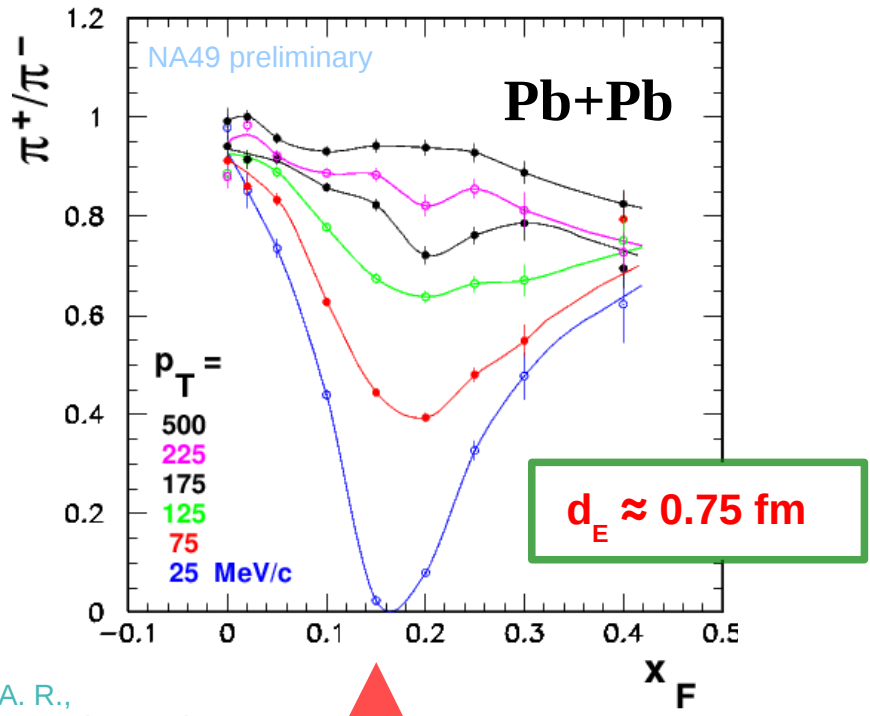
Repulsion (for  $\pi^+$ )  
Attraction (for  $\pi^-$ )

$$x_F = \frac{p_L}{p_L^{beam}}$$

(c.m.s.)

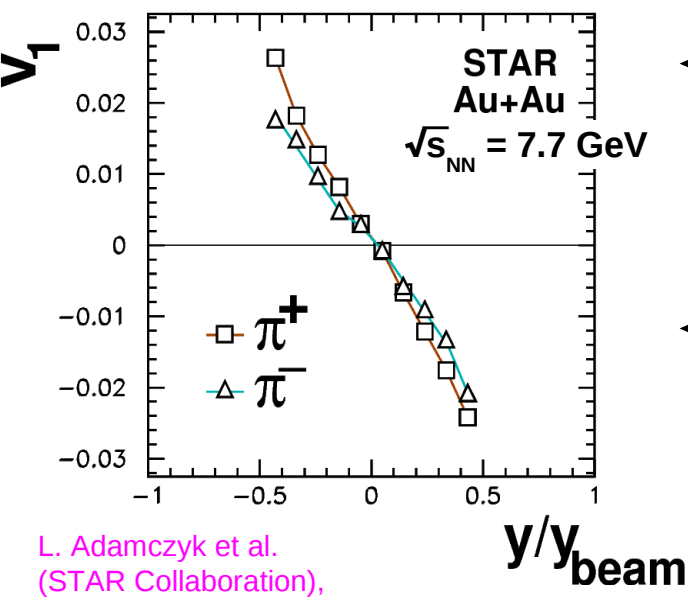
**Pb+Pb,  
peripheral**



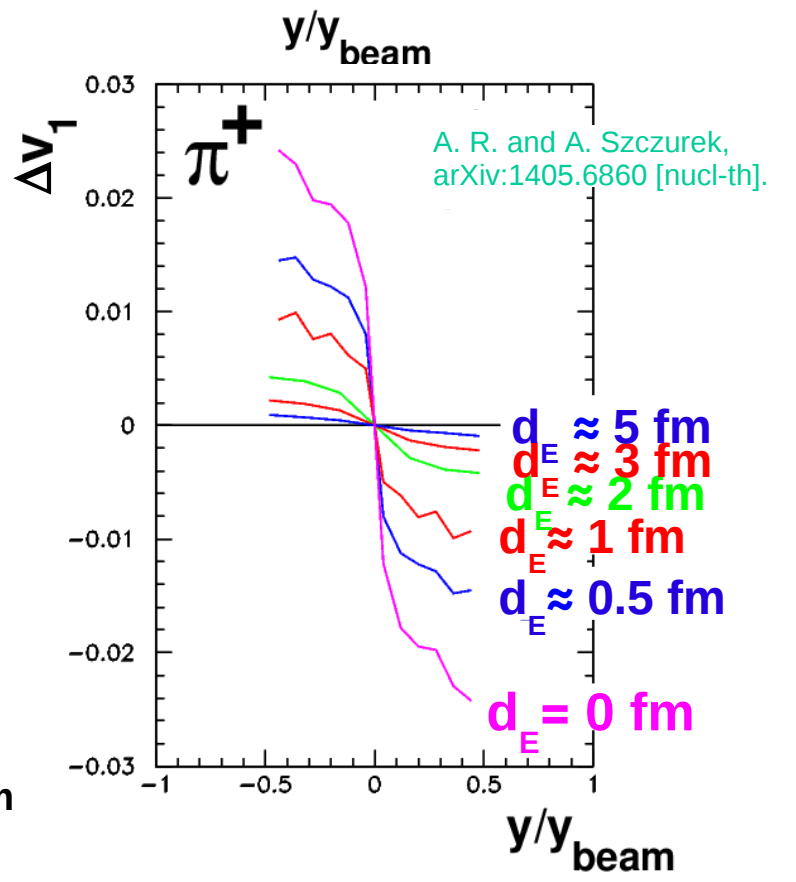
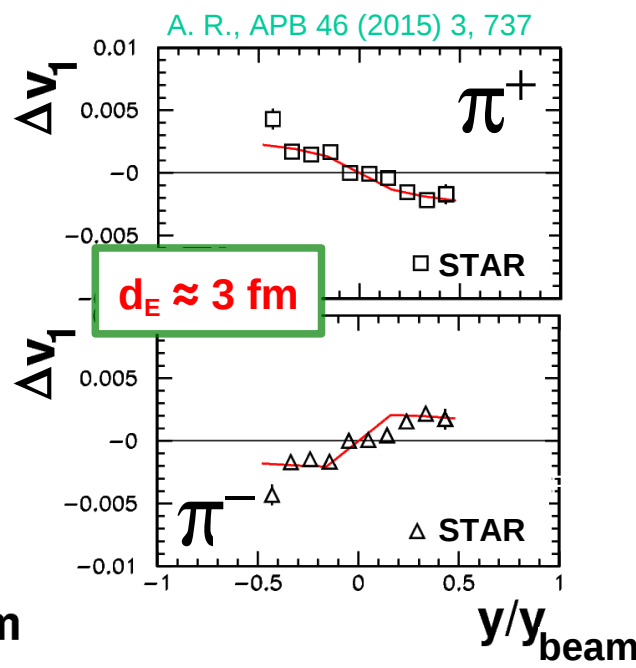


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$y = y_{\text{beam}}$



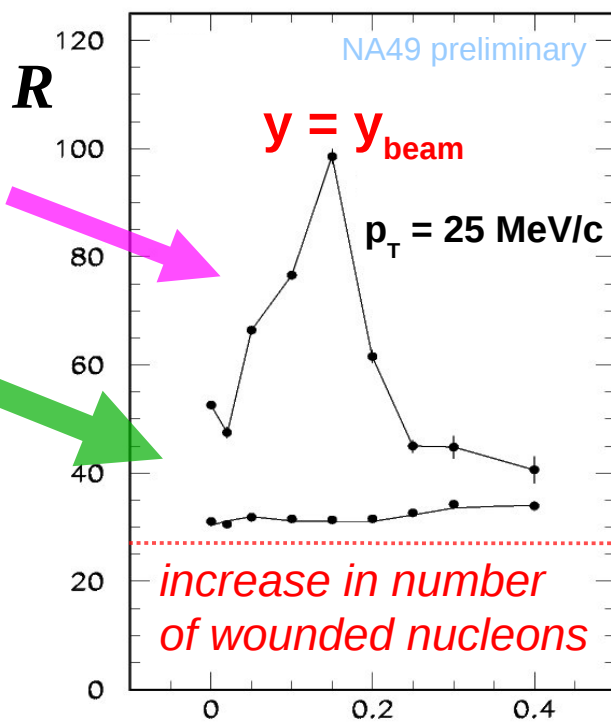
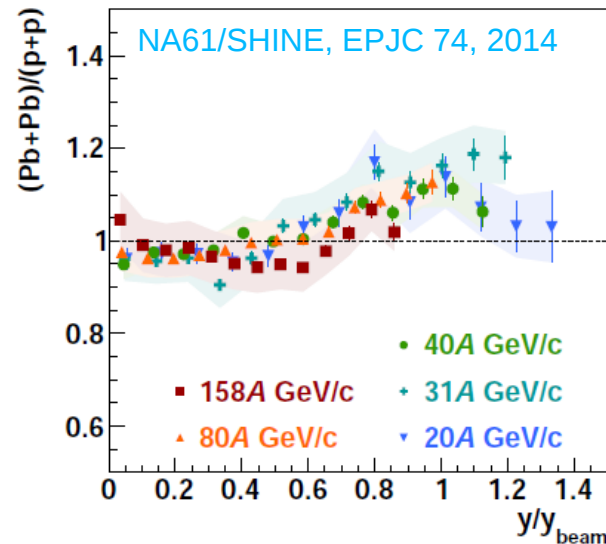
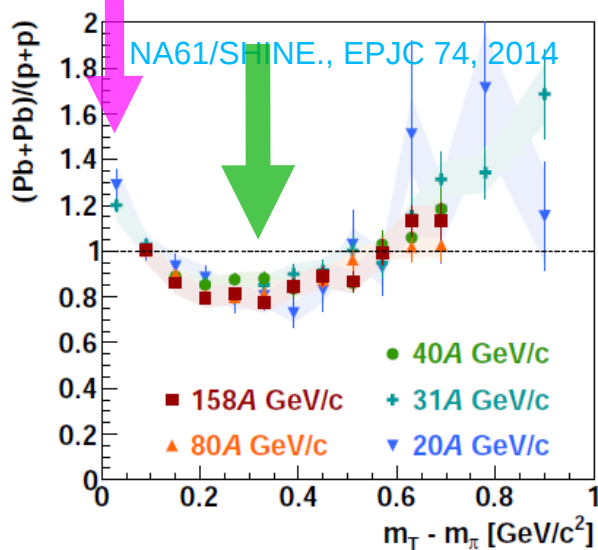
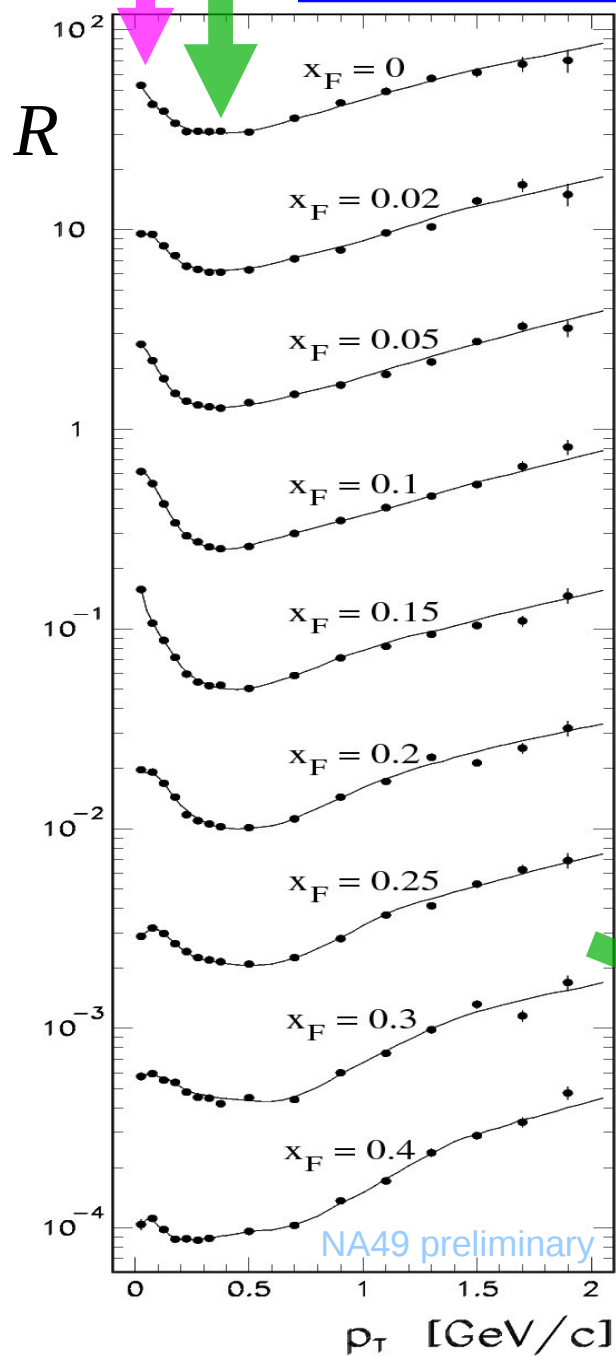
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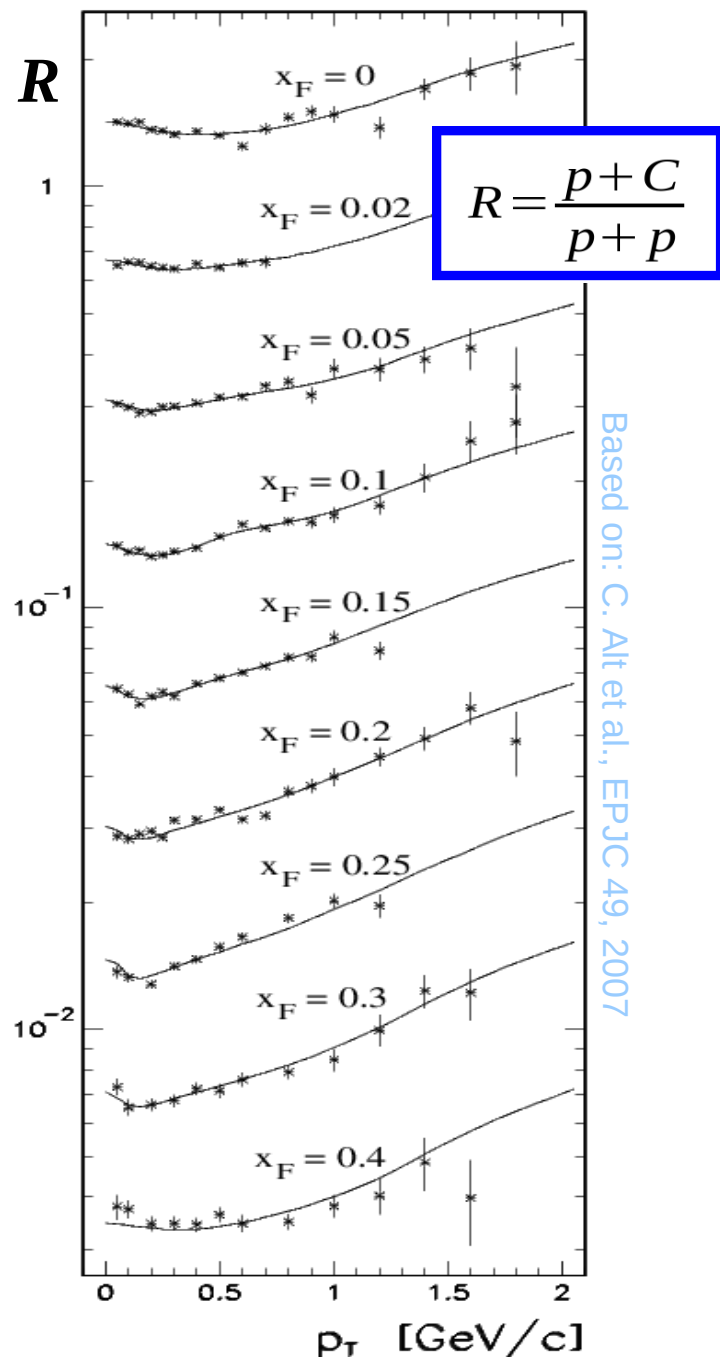
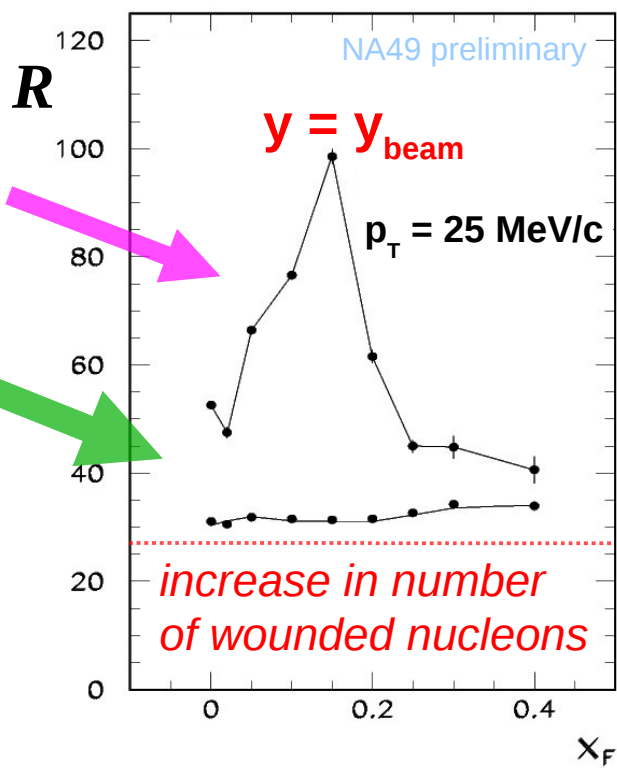
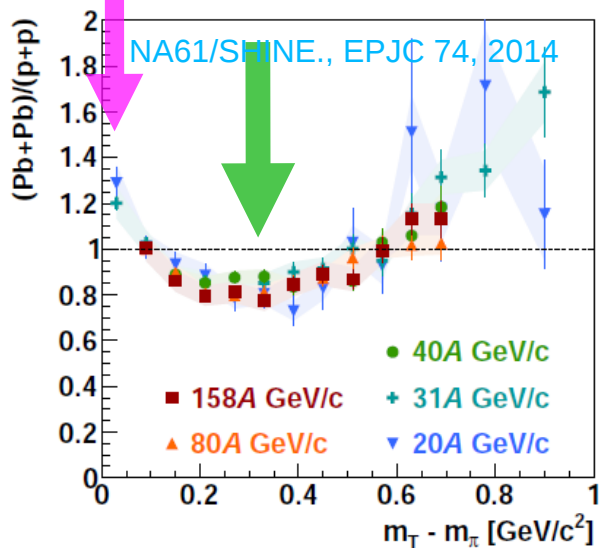
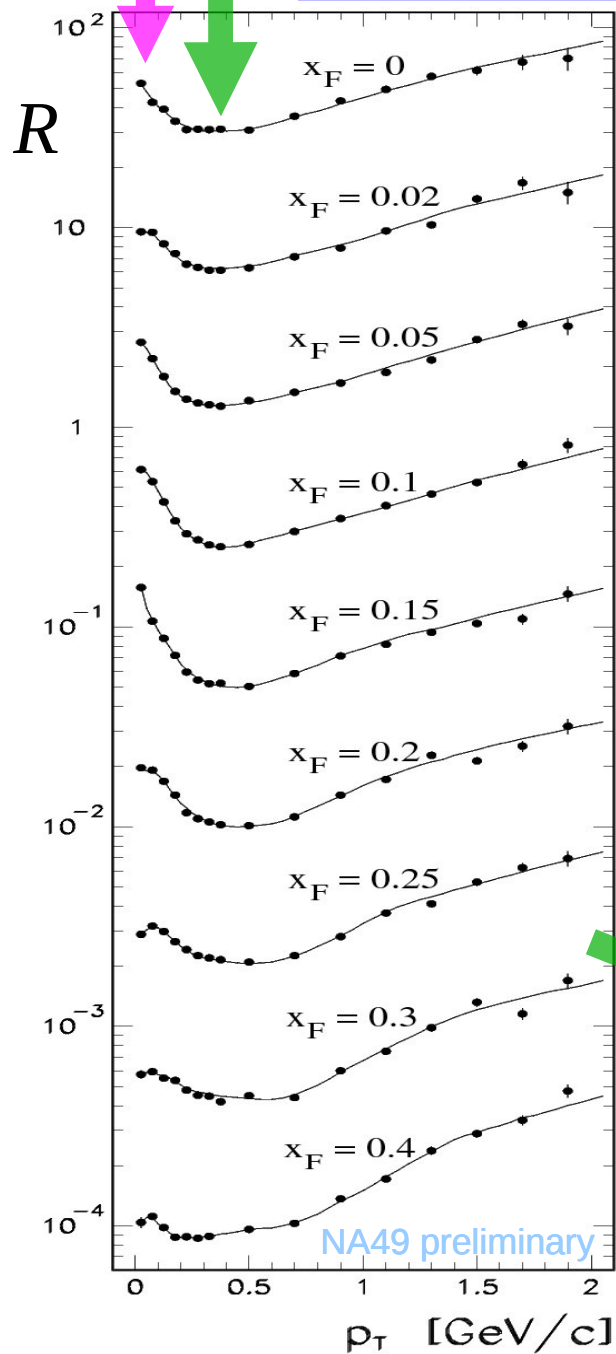
$$R = \frac{Pb+Pb}{p+p}$$

- NA49:  $(\pi^+ + \pi^-)/2$ ; Pb+Pb peripheral.
- NA61/SHINE:  $\pi^-$  only; Pb+Pb CENTRAL.



$$R = \frac{Pb + Pb}{p + p}$$

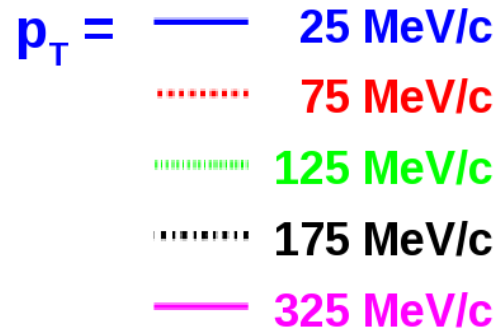
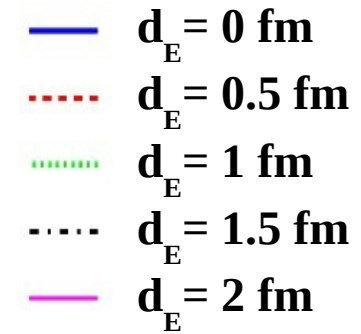
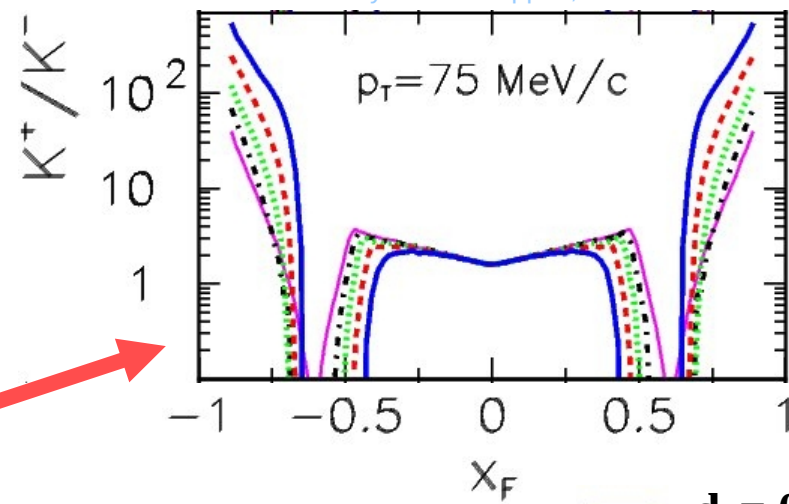
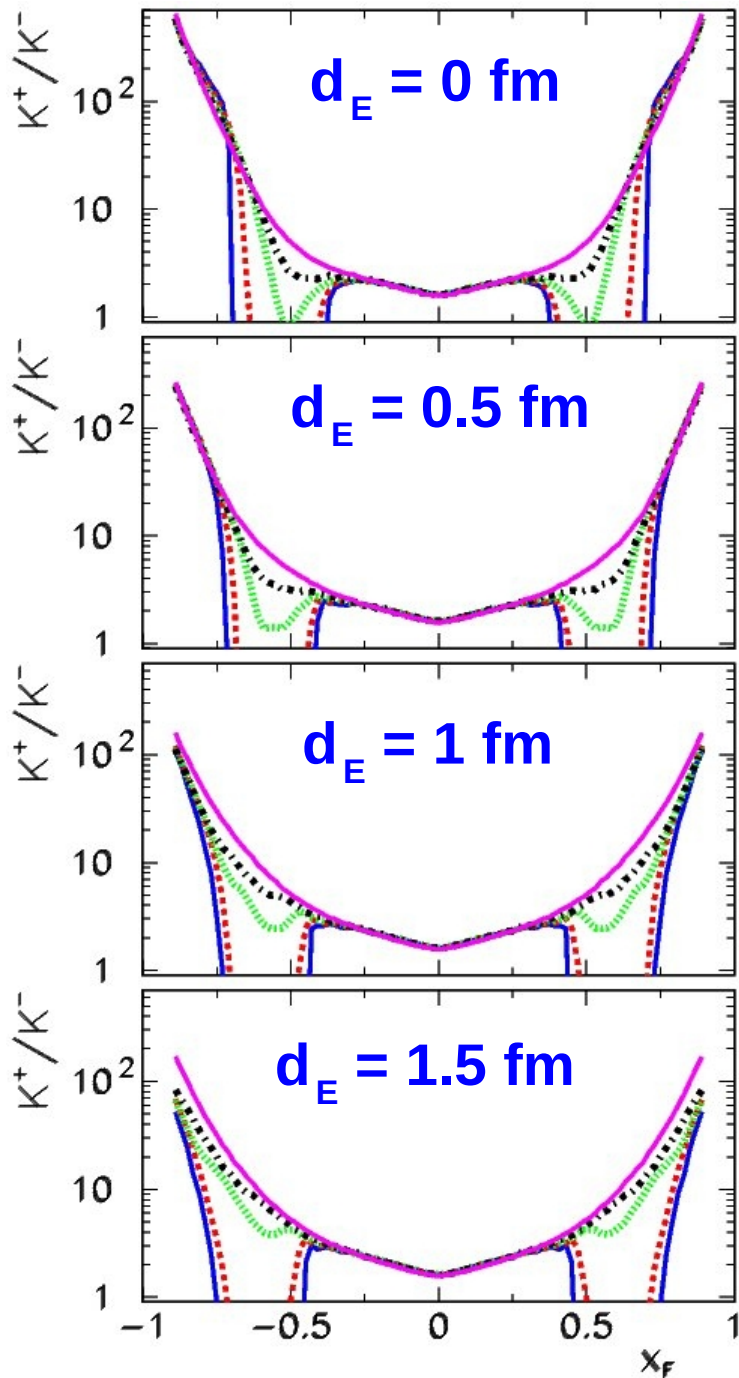
- NA49:  $(\pi^+ + \pi^-)/2$ ; Pb+Pb peripheral.
- NA61/SHINE:  $\pi^-$  only; Pb+Pb CENTRAL.



$$R = \frac{p + C}{p + p}$$

Based on: C. Alt et al., EPJC 49, 2007

A.R., Acta Phys. Pol. B42, 2011



- Large effect above  $x_F = 0.5$ .
- Dependence on initial conditions.
- **Very high  $x_F$**