Radiative pion capture in ²H, ³He and ³H



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Outline

- Introduction: elements of formalism
- Radiative pion capture on ²H, ³He and ³H
- Conclusions and outlook



Introduction

A very efficient momentum space framework to deal with nucleon-nucleon scattering, nucleon-deuteron scattering and many electroweak processes has been constructed and tested: Phys. Rept. 274, 107 (1996); Phys. Rept. 415, 89 (2005); Eur. Phys. J. A24, 31 (2005)

Limitations: nonrelativistic character and lack of the Coulomb force in the 3N

continuum

Calculations performed with semi-phenomenological 2N and 3N potentials: Bonn B, AV18, Nijmegen I and II, CD Bonn, Urbana IX, older chiral potentials from the Bonn/Bochum group and recently with the improved chiral potentials from E. Epelbaum *et al.*



Introduction

Methods developed originally for elastic and inelastic electron scattering, photodisintegration, and applied later to neutrino induced reactions and muon capture are now used to investigate the following processes

$$\pi^{-} + d \rightarrow \gamma + n + n$$

$$\pi^{-} + {}^{3}\text{H}e \rightarrow \gamma + {}^{3}\text{H}$$

$$\pi^{-} + {}^{3}\text{H}e \rightarrow \gamma + d + n$$

$$\pi^{-} + {}^{3}\text{H}e \rightarrow \gamma + p + n + n$$

$$\pi^{-} + {}^{3}\text{H} \rightarrow \gamma + n + n + n$$

These processes combine information from several areas (pion absorption ↔ pion photoproduction, weak processes, nuclear interactions) and should be ultimately studied within ChEFT



Introduction

General strategy in the few-nucleon systems:

- use (consistent) dynamical ingredients (2N and 3N potentials, electroweak current operators)
- solve the dynamical equations (Schrödinger equation, Lippmann-Schwinger equation, Faddeev equations)
- give predictions for nuclear structure and reaction observables
- confront results of theoretical calculations with experimental data to improve your input

Presented here results have been calculated with the AV18 2N and Urbana IX 3N potentials



Formalism





Formalism

Pion capture from the lowest K-shell of the pionic atom followed by photon emission studied with impulse approximation:

$$T_{fi} \propto \vec{\varepsilon} \cdot \left\langle \Psi_f \left| \vec{j}_A \right| \Psi_i \right\rangle$$

$$\vec{j}_A \approx \sum_{i=1}^A \tau_-(i) \vec{\sigma}(i)$$

nuclear axial current

final photon polarization vector



$$\psi_K(r) \equiv \psi_{100}(r) = \sqrt{\frac{(Z m' \alpha)^3}{\pi}} e^{-Z m' \alpha r}$$

$$m' \equiv \frac{m_{\pi}m_Z}{m_{\pi} + m_Z}$$
$$E_1 = -\frac{Z^2 \alpha^2 m'}{2}$$

reduced mass

negligible for Z=1,2 when compared to the pion or nucleon mass







$$\pi^- + d \rightarrow \gamma + n + n$$



$$\Gamma_{nn} = \frac{1}{2} \frac{1}{(2\pi)^2} \frac{2\pi\alpha}{f_\pi^2 M_\pi} \frac{(M_d'\alpha)^3}{\pi} \int_0^\pi d\theta_{p\gamma} \sin\theta_{p\gamma} \int_0^{2\pi} d\phi_{p\gamma} \int_0^{E_\gamma^{max,nn}} dE_\gamma E_\gamma \frac{1}{2} M_n p$$
$$\int_0^\pi d\theta_p \sin\theta_p \int_0^{2\pi} d\phi_p \frac{1}{3} \sum_{m_d} \sum_{m_1,m_2} \left(|N_{+1}(m_1, m_2, m_d)|^2 + |N_{-1}(m_1, m_2, m_d)|^2 \right)$$

best way to get Γ_{nn} from $d\Gamma_{nn}/dE_{\gamma}$



 $\pi^- + d \rightarrow \gamma + n + n$



 $\Gamma_{nn} = 0.318 \times 10^{15} \text{ 1/s (PW)}$ $\Gamma_{nn} = 0.328 \times 10^{15} \text{ 1/s (FSI)}$

similar results but completely different physics !



$$\pi^- + d \rightarrow \gamma + n + n$$

Earlier theoretical predictions for Γ_{nn} :

A. Reitan, Nucl. Phys. 87, 232 (1966): $3.32 \times 10^{14} \text{ 1/s} \rightarrow 4 \times 10^{14} \text{ 1/s}$ (corrected by ST)

M. Sotona and E. Truhlik, Nucl. Phys. A262, 400 (1976): 3.75×10^{14} 1/s (based on pion photoproduction data) 3.83×10^{14} 1/s (based on soft-pion limit+ corrections)

W. R. Gibbs, B. F. Gibson, and Q. J. Stephenson, Jr., Phys. Rev. C16, 327 (1977); 17, 856 (1978) (E) (4.2 ± 0.5) × 10¹⁴ 1/s

this contribution: 3.28×10^{14} 1/s



$$\pi^- + d \rightarrow \gamma + n + n$$

another way to get Γ_{nn} from neutron spectrum



$$\Gamma_{nn} = \frac{1}{2} \frac{1}{(2\pi)^2} \frac{2\pi\alpha}{f_\pi^2 M_\pi} \frac{(M'_d \alpha)^3}{\pi} \int_0^\pi d\theta_{p_\gamma} \sin \theta_{p_\gamma} \int_0^{2\pi} d\phi_{p_\gamma}$$
$$\int_0^\pi d\theta_{p_1} \sin \theta_{p_1} \int_0^{2\pi} d\phi_{p_1} \int_0^{E_1^{max}} dE_1 \frac{M_n^2 p_1 E_\gamma}{E_\gamma + M_n + p_1 \cos \theta_{\gamma 1}}$$
$$\frac{1}{3} \sum_{m_d} \sum_{m_1, m_2} \left(|N_{+1}(m_1, m_2, m_d)|^2 + |N_{-1}(m_1, m_2, m_d)|^2 \right)$$





neutron-neutron potential is changed by 1 % only in the ¹S₀ channel

neutron TOF spectra for $\theta_{\gamma 1} = 179^{\circ}$ normalized at QFS peak for s= 2.55 m

Radiative pion capture on ³He: triton channel

 $\pi^- + {}^{3}He \rightarrow \gamma + {}^{3}H$

 $\Gamma_{3H} = 2.059 \times 10^{15} \text{ 1/s} (2\text{NF})$ $\Gamma_{3H} = 2.132 \times 10^{15} \text{ 1/s} (2\text{NF}+3\text{NF})$

Radiative pion capture on ³He: triton channel

Earlier theoretical predictions for $\Gamma_{\rm 3H}$

- Fujii and D. Hall, Nucl. Phys. 32, 102 (1962).
 (8.32 → 4.28) × 10¹⁵ 1/s (corrected by Truöl 1974)
- P. Divakaran, Phys. Rev. 139, 3887 (1965).
 (0.97 → 3.88) × 10¹⁵ 1/s (corrected by Truöl 1974)
- D. Griffiths and C. Kim, Phys. Rev. 173, 1584(1968)
 2.32 × 10¹⁵ 1/s
- P. Pascual and A. Fujii, Nuovo Cimento 65, 411 (1970) (3.37 → 2.25) × 10¹⁵ 1/s (corrected by Truöl 1974)
- P. Truöl et al., Phys. Rev. Lett. 32, 1268 (1974) 3.60 × 10¹⁵ 1/s
- A. C. Phillips and F. Roig, Nucl. Phys. A234, 378 (1974) (3.1 - 3.7) × 10¹⁵ 1/s
- W. R. Gibbs et al., Phys. Rev. C18, 1761 (1978) 3.30 × 10¹⁵ 1/s

this contribution: 2.132×10^{15} 1/s

Radiative pion capture on ³He: two-body breakup $\pi^{-}+{}^{3}\text{H}e \rightarrow \gamma + d + n$

Radiative pion capture on ³He: three-body breakup

Radiative pion capture on ³He: three-body breakup

Radiative pion capture on ³He: comparison of twoand three-body breakup with best dynamics

Radiative pion capture on ³He: other predictions and data for breakup channels

 $\Gamma_{nd+nnp}/\Gamma_{3H}=1.2$

P. Truöl et al., Phys. Rev. Lett. 32, 1268 (1974) EXPERIMENT: (1.12 ± 0.05)

Radiative pion capture on ³H: three-neutron breakup

 \dot{p}_3

$$\pi^{-}+{}^{3}H \rightarrow \gamma + n + n + n$$

Radiative pion capture on ³H: three-neutron breakup

Radiative pion capture on ³H: other predictions and data

Fig. 4. Measured tritium spectrum with background subtracted. Curve a, Amado model and curve b, plane wave for 3n final state, from refs. ^{3, 21}), folded with the experimental resolution and corrected for spectrometer acceptance.

J. P. Miller et al., Nucl. Phys. A343, 347 (1980)

Calculations from A. C. Phillips and F. Roig, AIP Conf. Proc. No. 26, (1975)

 $\Gamma_{nnn} = 0.07 \times 10^{15} \, 1/s$

$$\Gamma_{\rm nnn} = 0.128 \times 10^{15} \, 1/s$$

Conclusions and outlook

- A very robust momentum space framework to deal with many electroweak processes has been applied to radiative pion capture processes
- First consistent results for ²H, ³He and ³H with realistic 2N and 3N potentials have been obtained in impulse approximation
- Sensitivity to properties of neutron-neutron interaction in the ²H case has been confirmed
- Comparisons with other theories yield a mixed picture
- Room for improvement: consistent 2N and 3N potentials as well as transition operators should be used for all radiative capture reactions
- New data are necessary to establish detailed relations among many processes (3N scattering, weak proton-proton capture, neutrino scattering, muon capture, pion absorption, pion photoproduction)

Investigations of double radiative pion capture are planned

Selected references:

W. K. H. Panofsky et al., Phys. Rev. 81 (1951) 565 K. M. Watson and R. N. Stuart, Phys. Rev. 82, 738 (1951) A. M. L. Messiah, Phys. Rev. 87, 639 (1952). N. M. Kroll and M. A. Ruderman, Phys. Rev. 93, 233 (1954) K. McVoy, Phys. Rev. 121, 1401 (1961) P. Truöl et al., Phys. Rev. Lett. 32, 1268 (1974) A. C. Phillips and F. Roig, Nucl. Phys. A 234, 378 (1974) J. A. Bistirlich et al., Phys. Rev. Lett. 36, 942 (1976) H. W. Baer, K. M. Crowe, and P. Truöl, Adv. Nucl. Phys. 9, 177 (1977) W. R. Gibbs, B. F. Gibson, and Q. J. Stephenson, Jr., Phys. Rev. C 11, 90 (1975); 12, 2130(E) (1975); Phys. Rev. C 16, 322 (1977); Phys. Rev. C 16, 327 (1977); 17, 856(E) (1978); Phys. Rev. C 18, 1761 (1978)

Selected references:

G. F. de Teramond, Phys. Rev C 16, 1976 (1977); 36, 691 (1987)
J. P. Miller et al., Nucl. Phys. A343, 341 (1980)
V. Bernard, N. Kaiser, and U.-G. Meiβner, Phys. Lett. B 383, 116, (1996)
H. W. Fearing et al., Phys. Rev. C 62, 054006 (2000)
A. Gårdestig and D. R. Phillips, Phys. Rev. C 73, 014002 (2006)
A. Gårdestig, Phys. Rev. C 74, 017001 (2006)
Q. Chen et al., Phys. Rev. C 77, 054002 (2008)
A. Gårdestig, J. Phys. G: Nucl. Phys. 36, 053001 (2009).

J. Golak et al., Phys. Rept. 415, 89 (2005)
J. Golak et al., Phys. Rev. C 90, 024001 (2014)
J. Golak et al., Phys. Rev. C 94, 034002 (2016)

Thank you !

Auxiliary slides

Formalism (cont.)

$$N^{\alpha} = \left\langle \Psi_{f \, m_f} \left| j^{\alpha} \right| \Psi_{i \, m_i} \right\rangle$$

from *ab initio* calculations in momentum space

Dynamical ingredients (1): 2N and 3N Hamiltonians

$$\begin{split} H_{2N} &= H_0^{2N} + V_{12} \\ H_{3N} &= H_0^{3N} + V_{23} + V_{13} + V_{12} + V_{123} \equiv H_0^{3N} + V_1 + V_2 + V_3 + V_4 \\ &\equiv H_0^{3N} + V_1 + V_2 + V_3 + \underbrace{V_4^{(1)} + V_4^{(2)} + V_4^{(3)}}_{V_4} \end{split}$$

used to generate nuclear bound and scattering states contain 2N and 3N potentials

Formalism (cont.)

Dynamical ingredients (2): nuclear single-nucleon, 2N and 3N current operators

$$\begin{split} j_{2N} &= j_1 + j_2 + j_{12} \\ j_{3N} &= j_1 + j_2 + j_3 + j_{12} + j_{23} + j_{13} + j_{123} \\ &\equiv j_1 + j_{23} + j_2 + j_{13} + j_3 + j_{12} + \underbrace{j_{123}^{(1)} + j_{123}^{(2)} + j_{123}^{(3)}}_{j_{123}} \\ &\equiv \underbrace{j_1 + j_{23} + j_{123}^{(1)}}_{j(1)} + \underbrace{j_2 + j_{13} + j_{123}^{(2)}}_{j(2)} + \underbrace{j_3 + j_{12} + j_{123}^{(3)}}_{j(3)} \end{split}$$

describe interactions of an external probe with a nuclear system

Formalism (reactions with ²H)

$$\begin{split} H_{2N} | \psi_d \rangle &= E_d | \psi_d \rangle & \text{deuteron state with } E_d < 0 \\ N^{\alpha} &\equiv \left\langle \psi'_d \right| j^{\alpha} | \psi_d \rangle & \text{elastic channel} \\ N^{\alpha} &\equiv \left\langle \psi^{(-)} \right| j^{\alpha}_{2N} | \psi_d \rangle &= {}_a \left\langle \vec{p}_o \right| \left(1 + t_{12} G_0^{2N}\right) j^{\alpha}_{2N} | \psi_d \rangle & \text{break-up channel} \\ H_{2N} | \psi^{(-)} \rangle &= E | \psi^{(-)} \rangle, \quad E = \frac{p_0^2}{m} > 0 \\ t_{12} &= V_{12} + t_{12} G_0^{2N} (E + i\varepsilon) V_{12} & \text{Lippmann-Schwinger equation} \\ G_0^{2N} (E) &\equiv \lim_{\varepsilon \to 0^+} \frac{1}{E + i\varepsilon - H_0^{2N}} & \text{free 2N propagator} \end{split}$$

$$H_{3N} |\Psi\rangle = E_b |\Psi\rangle$$

3N bound state with $E_b < 0$ generated by the Faddeev equation

$$N^{\lambda} = \left\langle \Psi' \middle| j_{3N}^{\lambda} \middle| \Psi \right\rangle$$

elastic or quasielastic channel with initial and final bound states

$$N^{\lambda} = \left\langle \Psi_{f}^{(-)} \left| j_{3N}^{\lambda} \right| \Psi_{i} \right\rangle$$

two-body or three-body break-up channel with final scattering states

$$\left|\Psi_{f}^{(-)}\right\rangle = \lim_{\varepsilon \to 0^{+}} \frac{-i\varepsilon}{E - i\varepsilon - H_{3N}} \left|\phi_{f}\right\rangle$$

formal definition including the channel state

Operators in 3N space:

(1) 3N force decomposed as $V_4 = V_4^{(1)} + V_4^{(2)} + V_4^{(3)}$

 $V_4^{(i)}$ is symmetric under the exchange of nucleons j and k, $i \neq j \neq k \neq i$

(2) free 3N propagator

$$G_0^{3N}(E) \equiv \lim_{\varepsilon \to 0^+} \frac{1}{E + i\varepsilon - H_0^{3N}}$$
(3) 2N off-shell t-matrix generated via LSE:

$$t_1 = V_1 + V_1 G_0^{3N} t_1$$

(4) permutation operator:

 $P = P_{12}P_{23} + P_{13}P_{23}$

1

Auxiliary equation for $|U^{\lambda}\rangle \equiv |U(j^{\lambda}, E_{c.m.}, Q)\rangle$

3N internal energy

magnitude of the three momentum transfer

$$\begin{aligned} \left| U^{\lambda} \right\rangle &= \left\{ t_1 \, G_0^{3N} + \frac{1}{2} \left(1 + P \right) V_4^{(1)} \, G_0^{3N} \left(1 + t_1 \, G_0^{3N} \right) \right\} \left(1 + P \right) j^{\lambda}(1) \left| \Psi_i \right\rangle \\ &+ \left\{ t_1 \, G_0^{3N} \, P + \frac{1}{2} \left(1 + P \right) V_4^{(1)} \, G_0^{3N} \left(1 + t_1 \, G_0^{3N} \right) P \right\} \left| U^{\lambda} \right\rangle \end{aligned}$$

Quadratures

$$N_{Nd}^{\lambda} = \left\langle \phi_{Nd} \left| \left(1 + P \right) j^{\lambda}(1) \right| \Psi_{i} \right\rangle + \left\langle \phi_{Nd} \left| P \right| U^{\lambda} \right\rangle$$
$$N_{3N}^{\lambda} = \left\langle \phi_{3N} \left| \left(1 + P \right) j^{\lambda}(1) \right| \Psi_{i} \right\rangle + \left\langle \phi_{3N} \left| t_{1} G_{0}^{3N} \left(1 + P \right) j^{\lambda}(1) \right| \Psi_{i} \right\rangle$$
$$+ \left\langle \phi_{3N} \left| P \right| U^{\lambda} \right\rangle + \left\langle \phi_{3N} \left| t_{1} G_{0}^{3N} P \right| U^{\lambda} \right\rangle$$

to obtain nuclear matrix elements for arbitrary exclusive kinematics ! Semi-exclusive and inclusive observables are calculated by suitable integrations over the phase space domains.

